

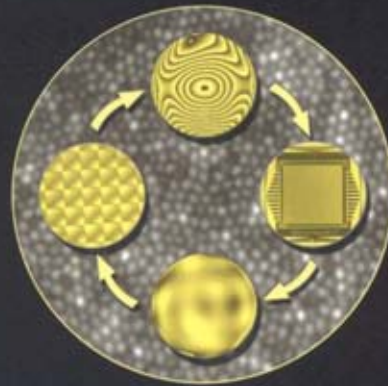
AO System Design for Vision Science

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Indiana University
School of Optometry



Adaptive Optics for Vision Science

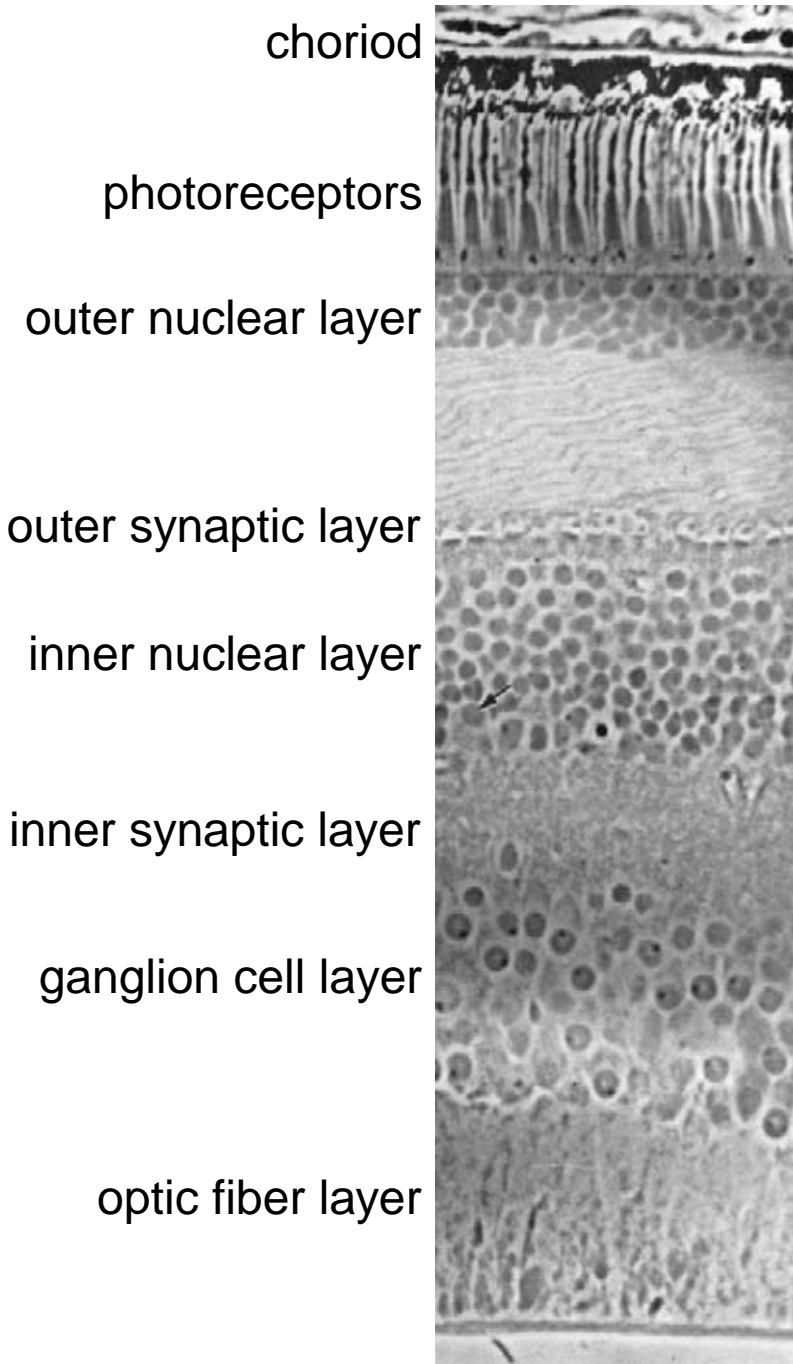


EDITED BY

JASON PORTER,
HOPE QUEENER, JULIANNA LIN,
KAREN THORN, and ABDUL AWWAL

Motivation for High-Resolution Retinal Imaging

- 1. Vision begins at the level of the cell.**
 - **Light collecting properties**
 - **Development of the normal eye.**
(Track retinal processes that vary over time.)
 - **Functionality of the various retinal layers.**
- 2. Pathogenesis begins at the level of the cell.**
 - **More precise and timely diagnosis**
 - **More effective monitoring**
 - **Better informed treatment**



Cell Populations

5 million pigment epithelium

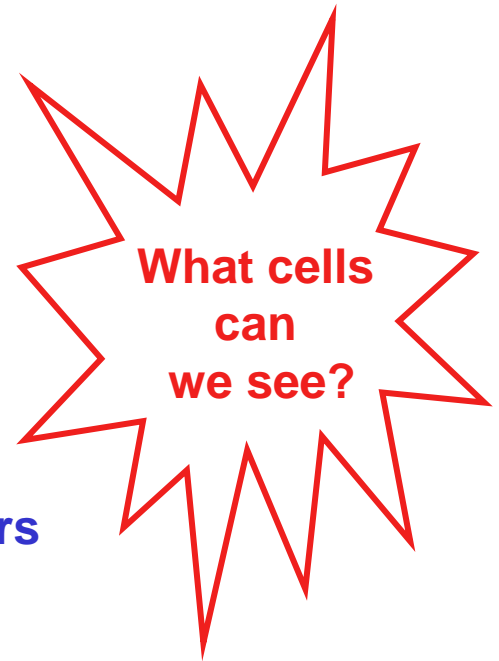
{ 6 million cones
117 million rods

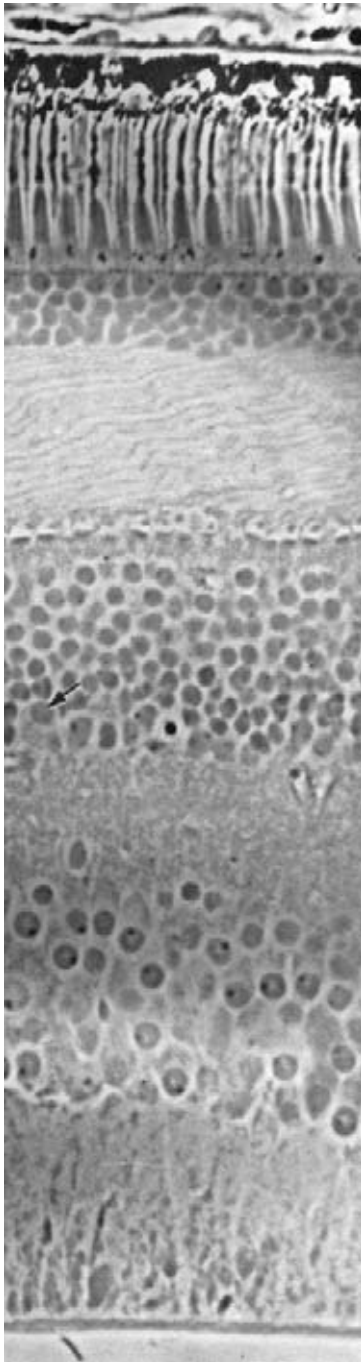
{ ? million bipolar
? million amacrine
? million horizontal
? million Muller

1 million ganglion

1 million nerve fibers

100 μm





Camera parameters important for imaging single cells in the living human retina

1. Lateral resolution
2. Axial resolution
3. Exposure time
4. Sensitivity
5. Contrast enhancement

Point Spread Function vs. Pupil Size

1 mm

2 mm

3 mm

4 mm

5 mm

6 mm

7 mm

Perfect Eye

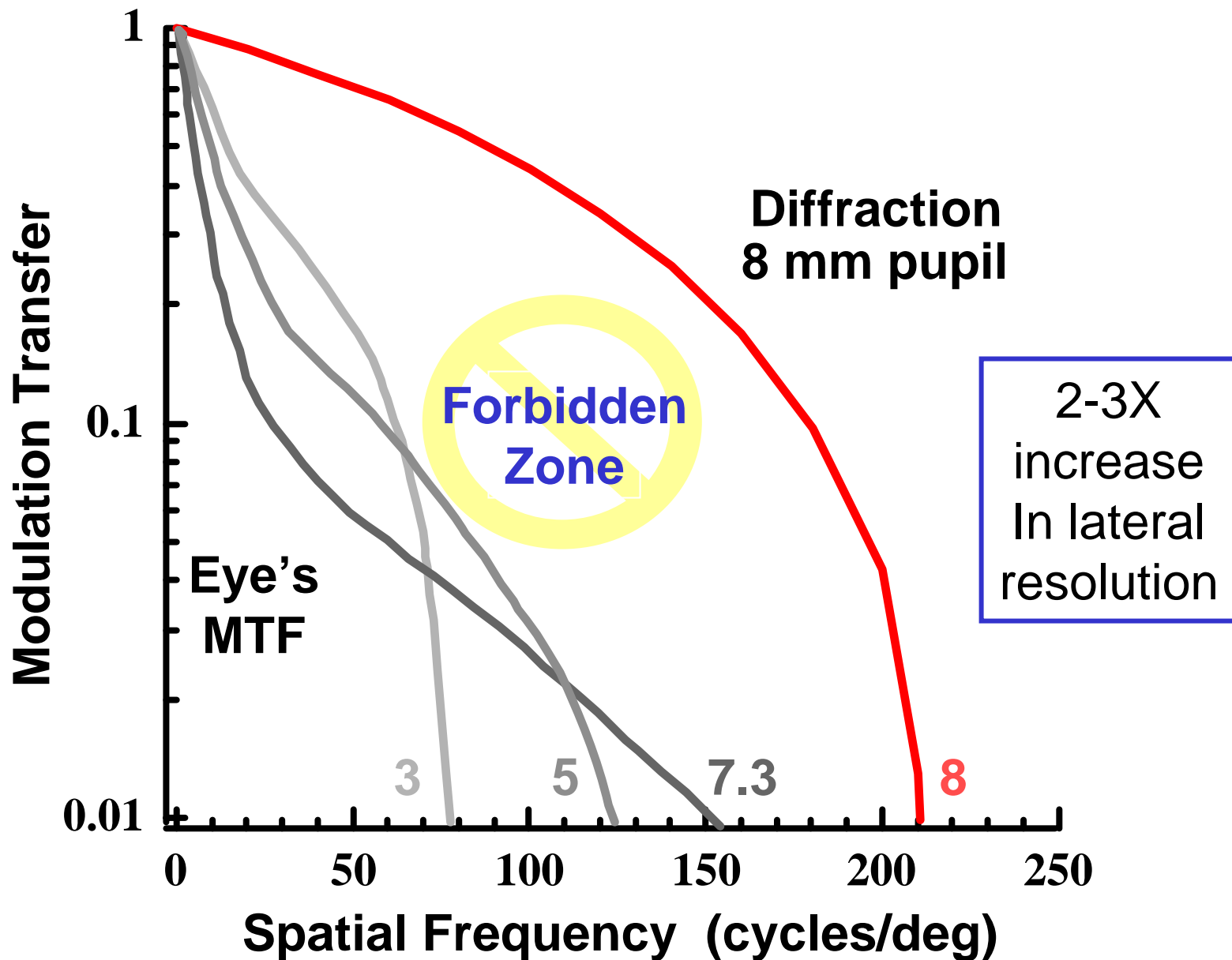
Typical Eye

AO

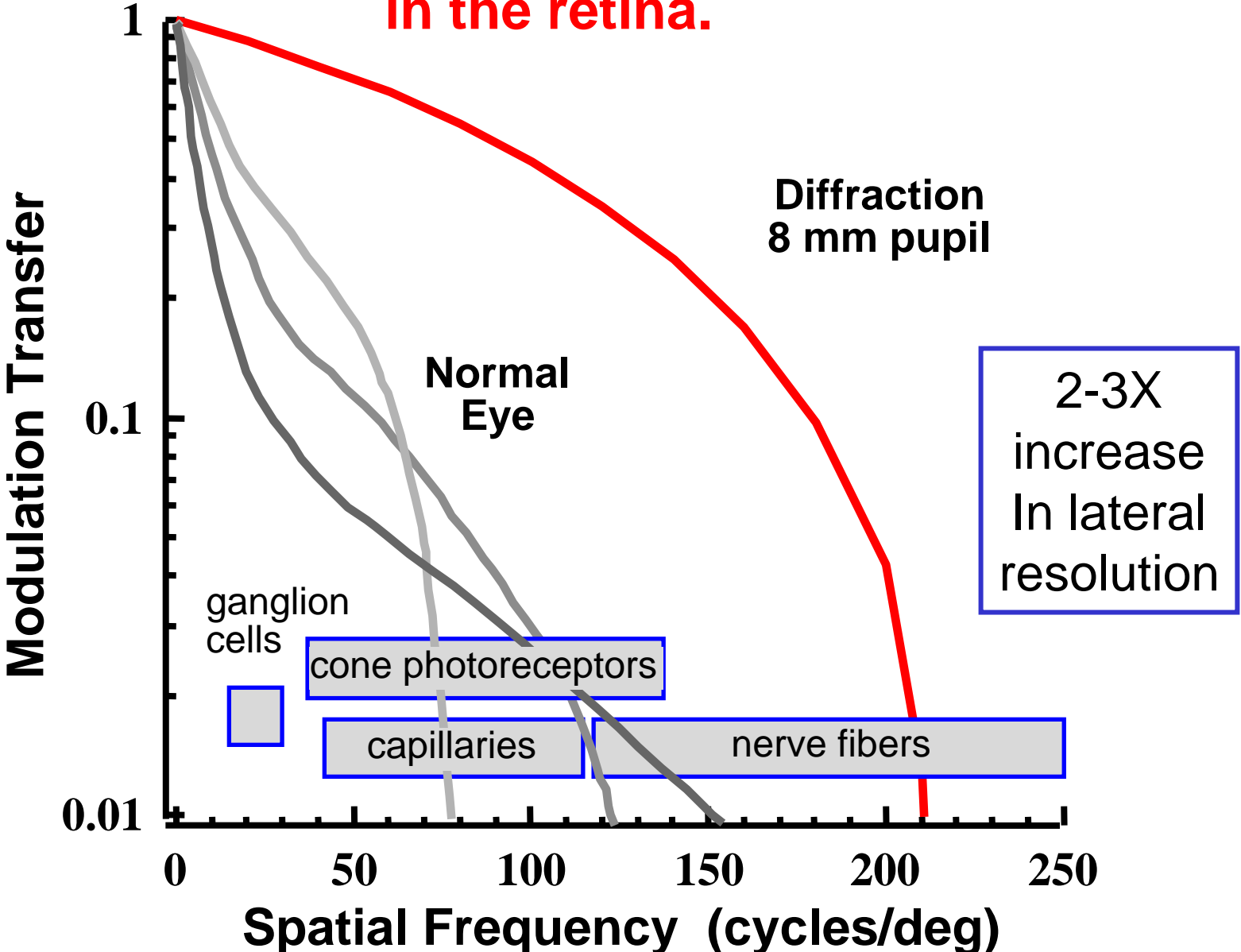
Courtesy A. Roorda



Aberrations of the eye limit access to high contrast and high spatial frequencies.



Aberration correction across a large pupil substantially increases the contrast and spatial detail of single cells in the retina.



Adaptive Optics In Astronomy:

- Proposed in 1953 (Horace Babcock)
- First Implemented in mid-1970s



The laser emerging from the dome of the 120" Shane Telescope at the Lick Observatory is used for measuring atmospheric aberrations

First Use of a Deformable Mirror In the Eye:

A. W. Dreher, J. F. Bille, R. N. Weinreb, "Active optical depth resolution improvement of the laser tomographic scanner," *Appl. Opt.* **24**, 804-808 (1989).

First use of wavefront sensing to correct higher order aberrations

First demonstration of retinal imaging and vision improvement by correcting higher order aberrations

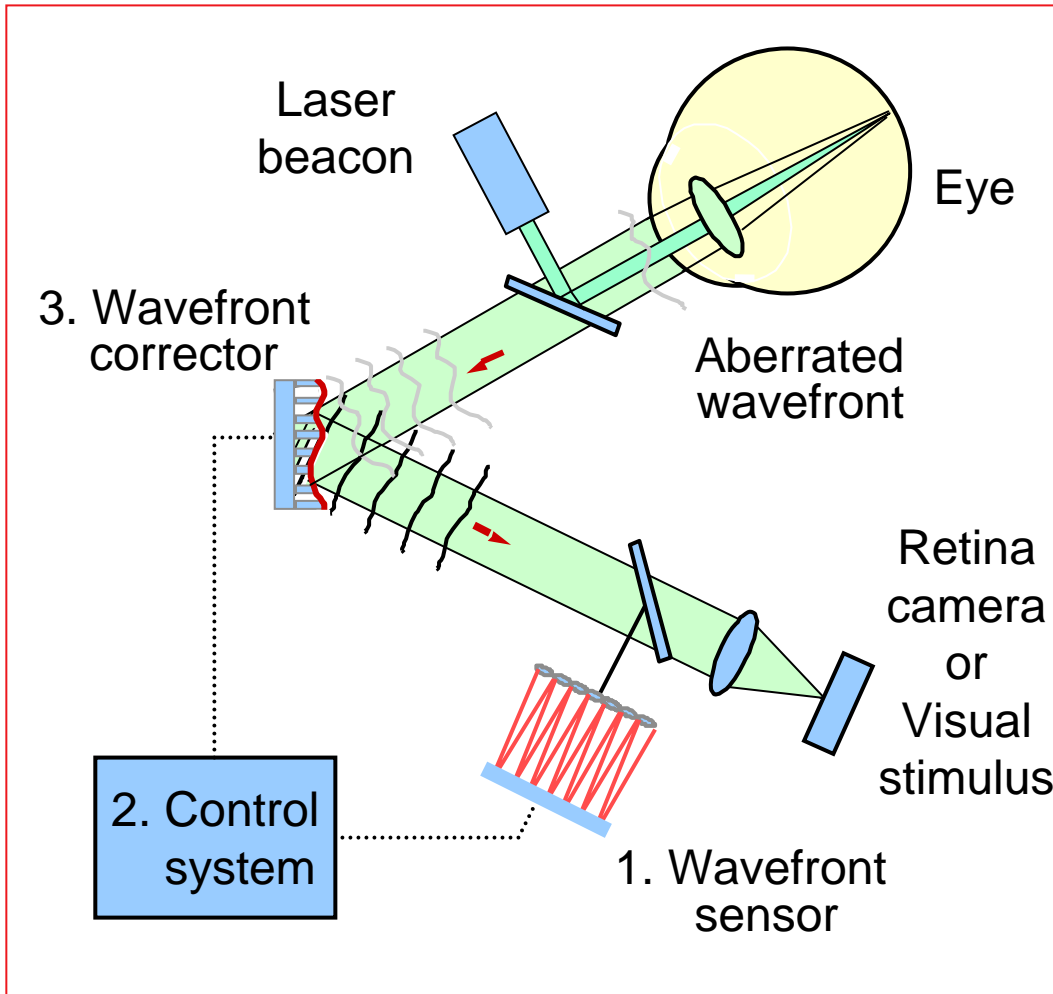


J. Liang, D.R. Williams, and D.T. Miller, "Supernormal vision and high resolution retinal imaging through adaptive optics,"
J. Opt.Soc. Am. A., 14, 2884-2892 (1997).

First AO images on a cone dystrophy patient.
Demonstrated that the mosaic was patchy due
to photoreceptor dropout

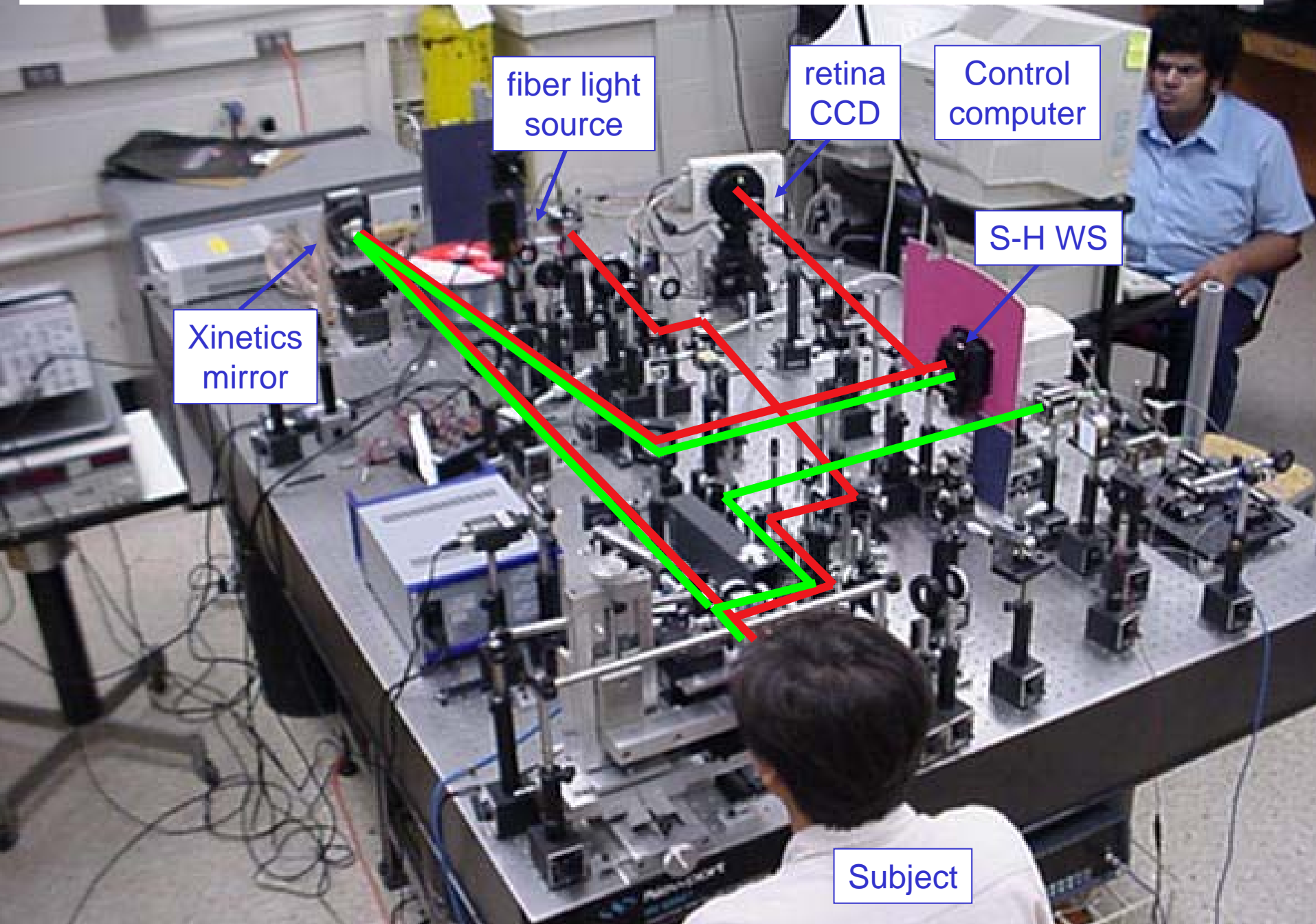
A. Roorda, "Adaptive optics ophthalmoscopy,"
J Ref Surgery. 2000;16:S602-607.

Adaptive optics measures and corrects the wave aberrations of the eye



- Permits large pupil (diffraction & NA)
- increased lateral resolution
- Increased collection efficiency of reflected light
- increased axial resolution (SLO)

Flood-illuminated AO retina camera



Numerous types of wavefront sensors are available

Common path interferometer

Phasing shifting interferometer

Shack-Hartmann sensor

Shearing interferometer

Pyramid sensor

Curvature sensor

Phase diversity

etc.

Noise sources in the Shack-Hartmann wavefront sensor

1. Noise sources: **photon noise, CCD read noise**

	<u>photons/lenslet</u>
astronomy:	~100 to close loop
vision:	>500,000

2. Spatial sampling error: **undetected aberrations (governed by lenslet dia.)**

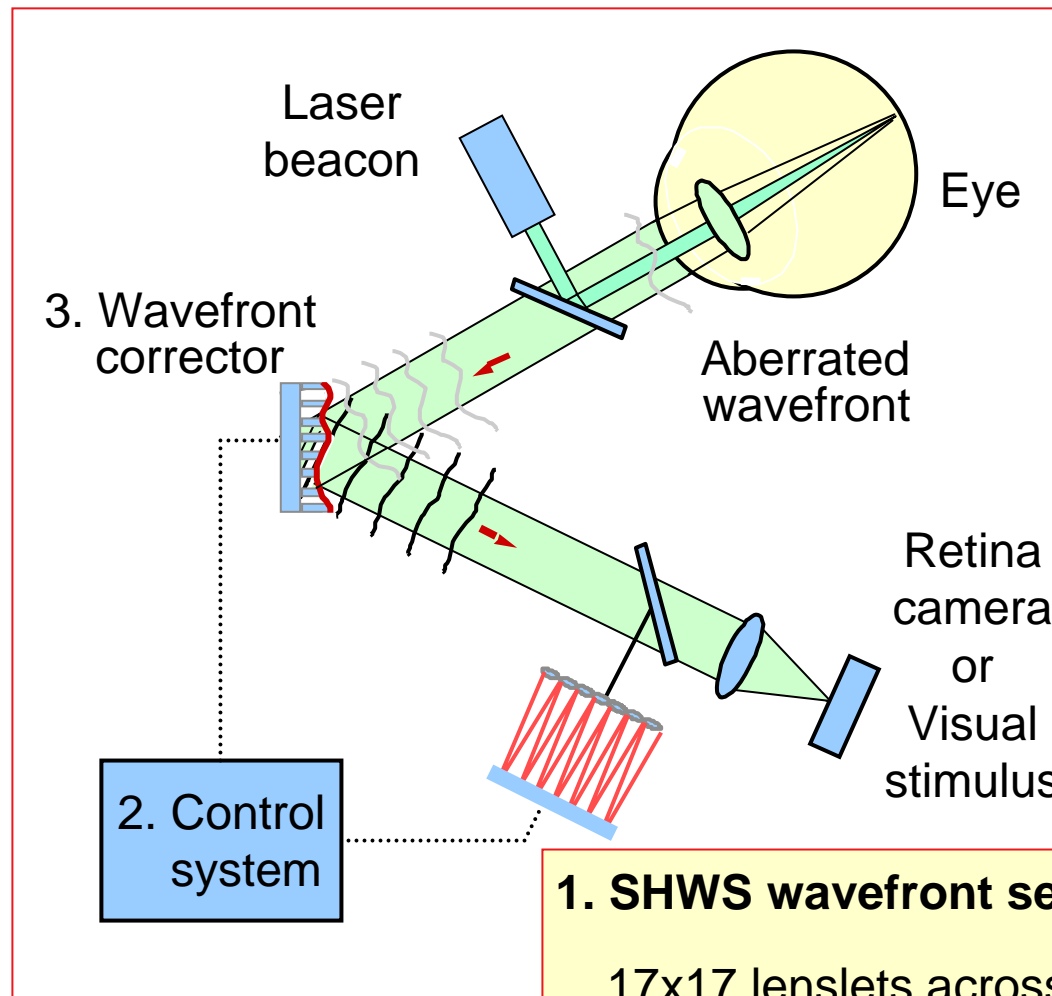
	<u>lenslet dia. and spacing</u>
astronomy:	~ r_o , uniform sampling
vision:	?, ?

3. Temporal sampling error: **undetected aberrations (1-6 Hz) (governed by frame rate & exposure)**

Typical SH sensor parameters for the eye

1. # of lenslets: ~200 (>1/2 million photons/lenslet for $8 \mu\text{W}$ entering eye.)
2. Lenslet array: lenslet diameter = $400 \mu\text{m}$
(17x17 for 6.8 mm pupil)
focal length = 24mm
3. # pixels across dot core 5 to 14
4. Wavelength 633 to 850 nm

Indiana adaptive optics retina camera



1. SHWS wavefront sensor

17x17 lenslets across 6.8 mm pupil
~7 pix across each focal spot
 $\lambda = 0.78 \mu\text{m}$ SLD beacon
6 μW enters eye
Inexpensive areal CCD



OKO membrane mirror



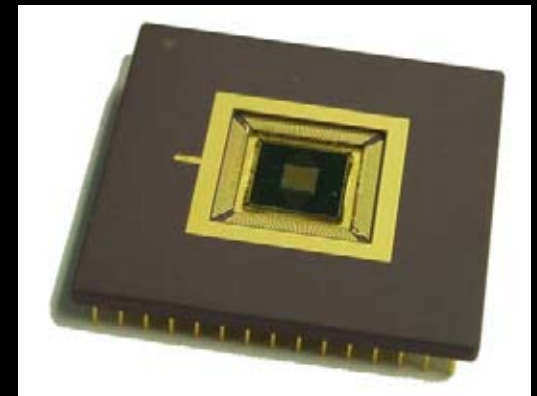
AOptix bimorph
"fast, simple, & robust"



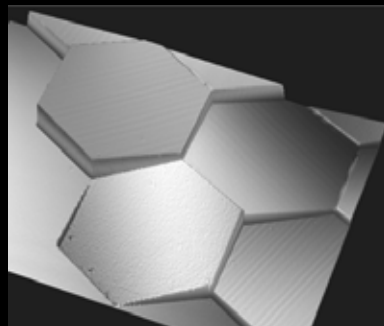
Hamamatsu LC-SLM



Xinetics deformable
mirror



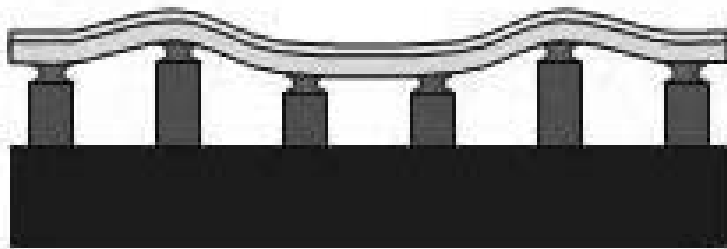
Boston
Micromachines
Corp.



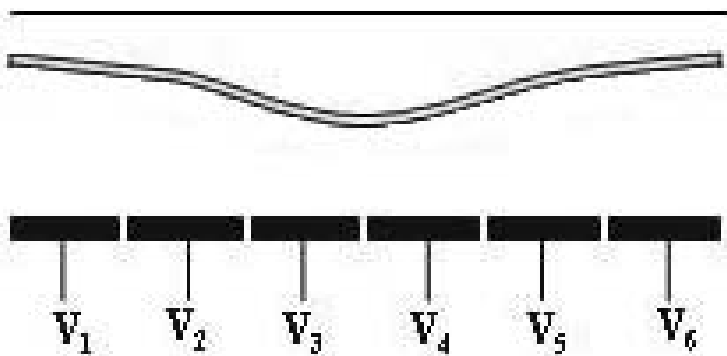
Iris AO

Types of wavefront correctors

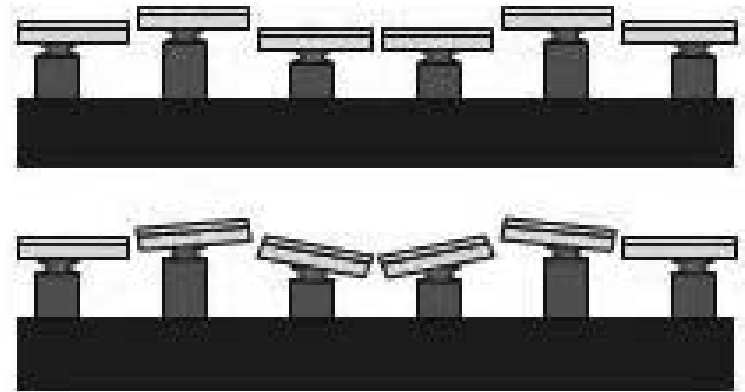
(Which one is right for my application?)



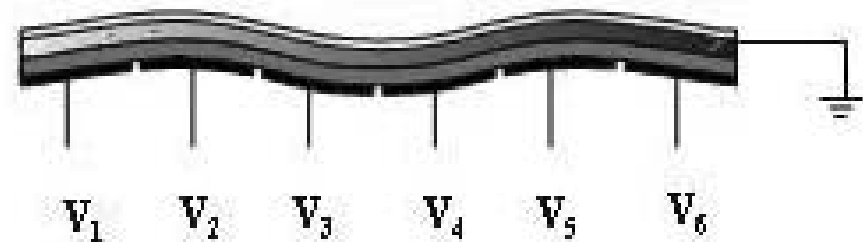
a) Discrete Actuator



c) Membrane



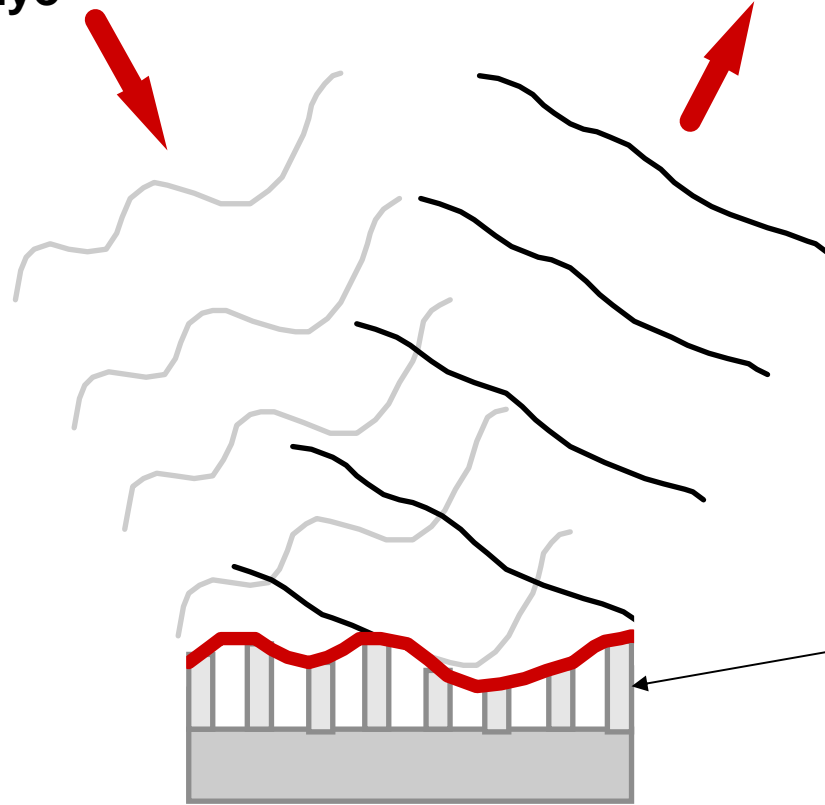
b) Segmented: Piston only
and Piston/tip/tilt



d) Bimorph

Wavefront Corrector Performance

Input from
Eye

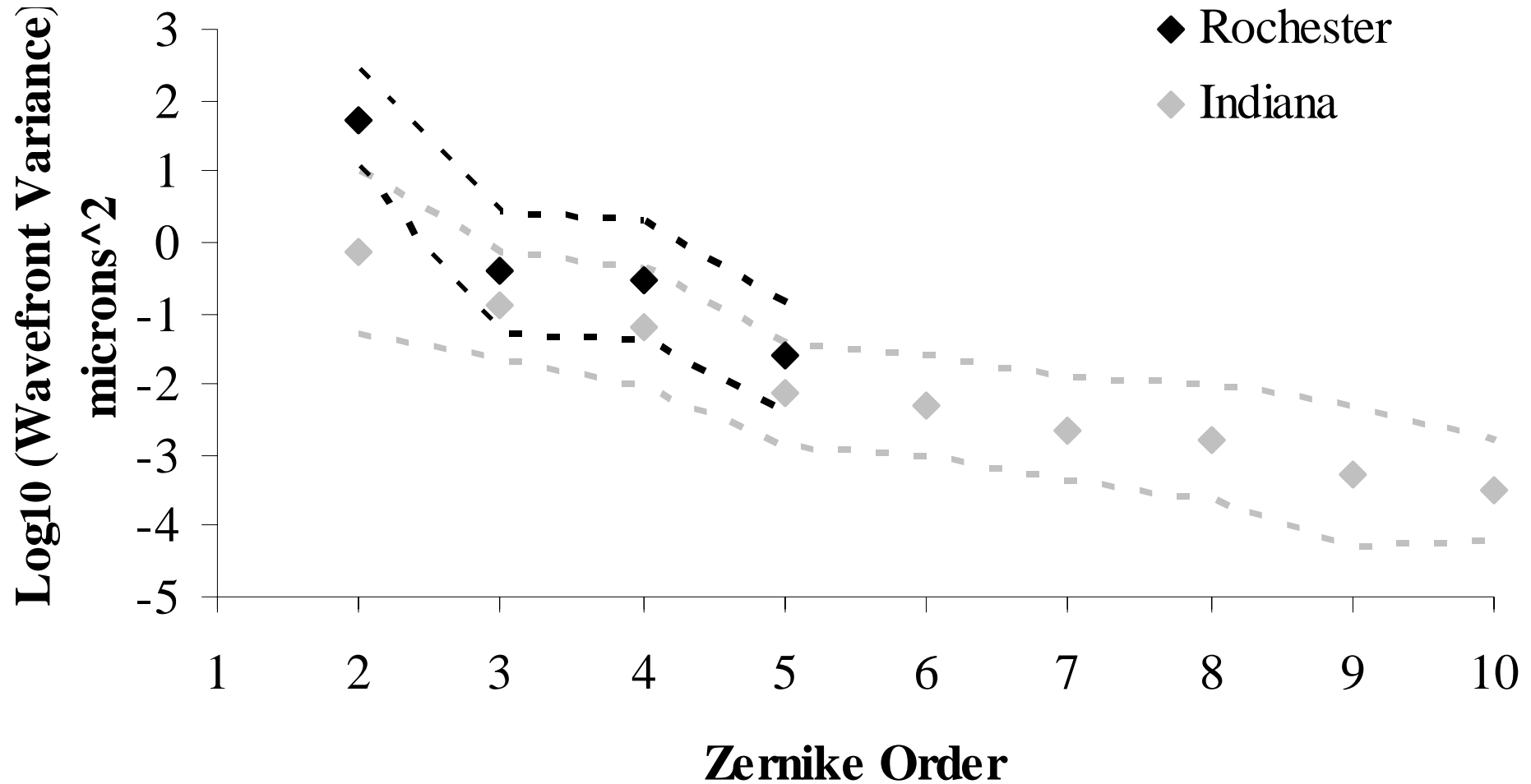


actuators
push & pull
on mirror
surface

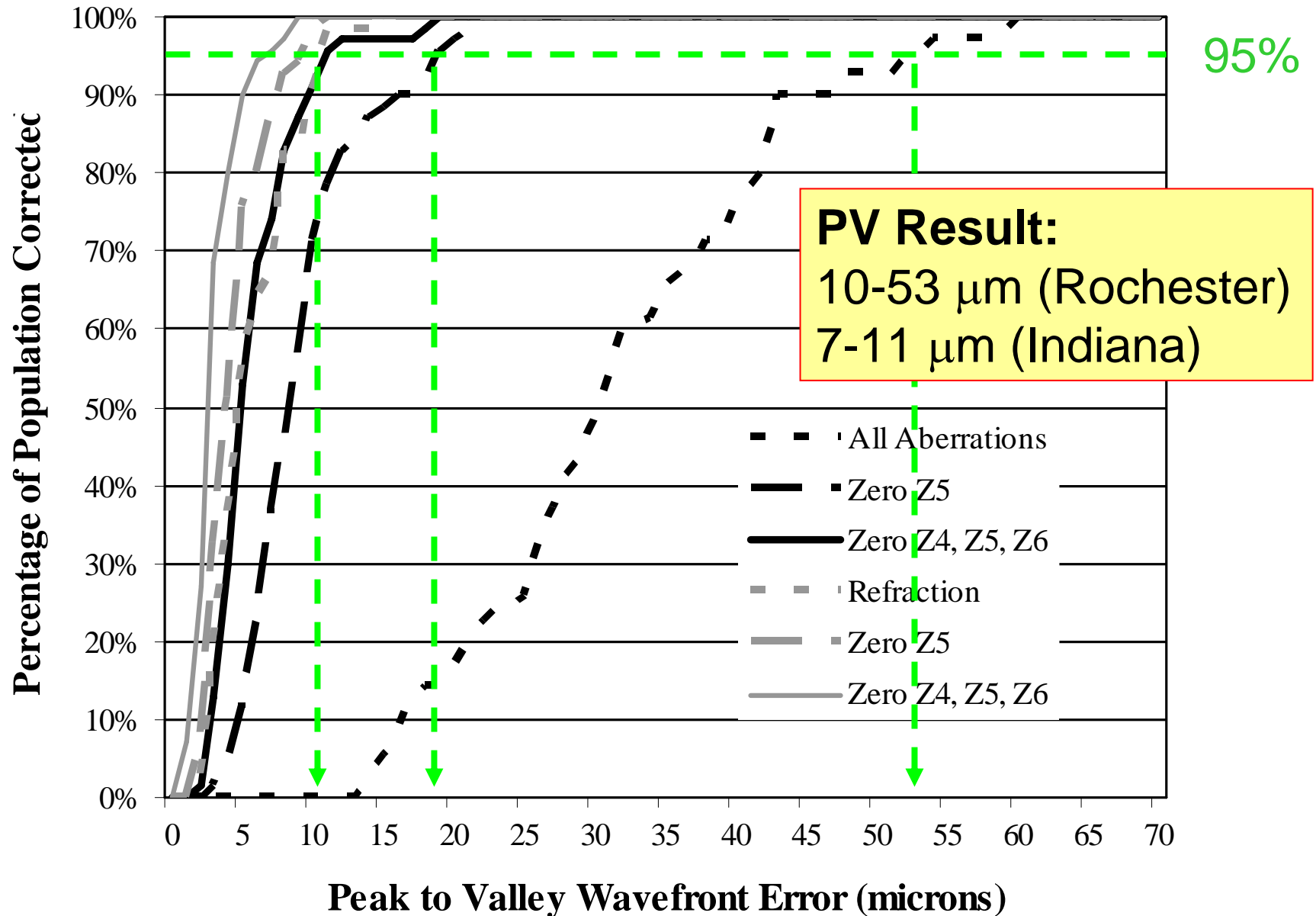
**Most important
parameters:**

- Actuator stroke
- Actuator number
- Actuator influence function
- Speed
- Reflectivity
- Diameter
- Cost!

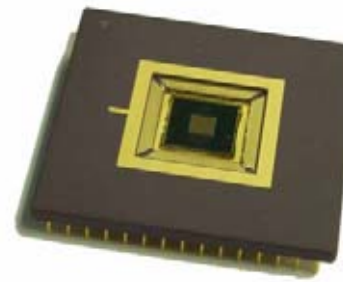
Aberrations in two populations of 70 normal eyes for 7.5 mm pupil



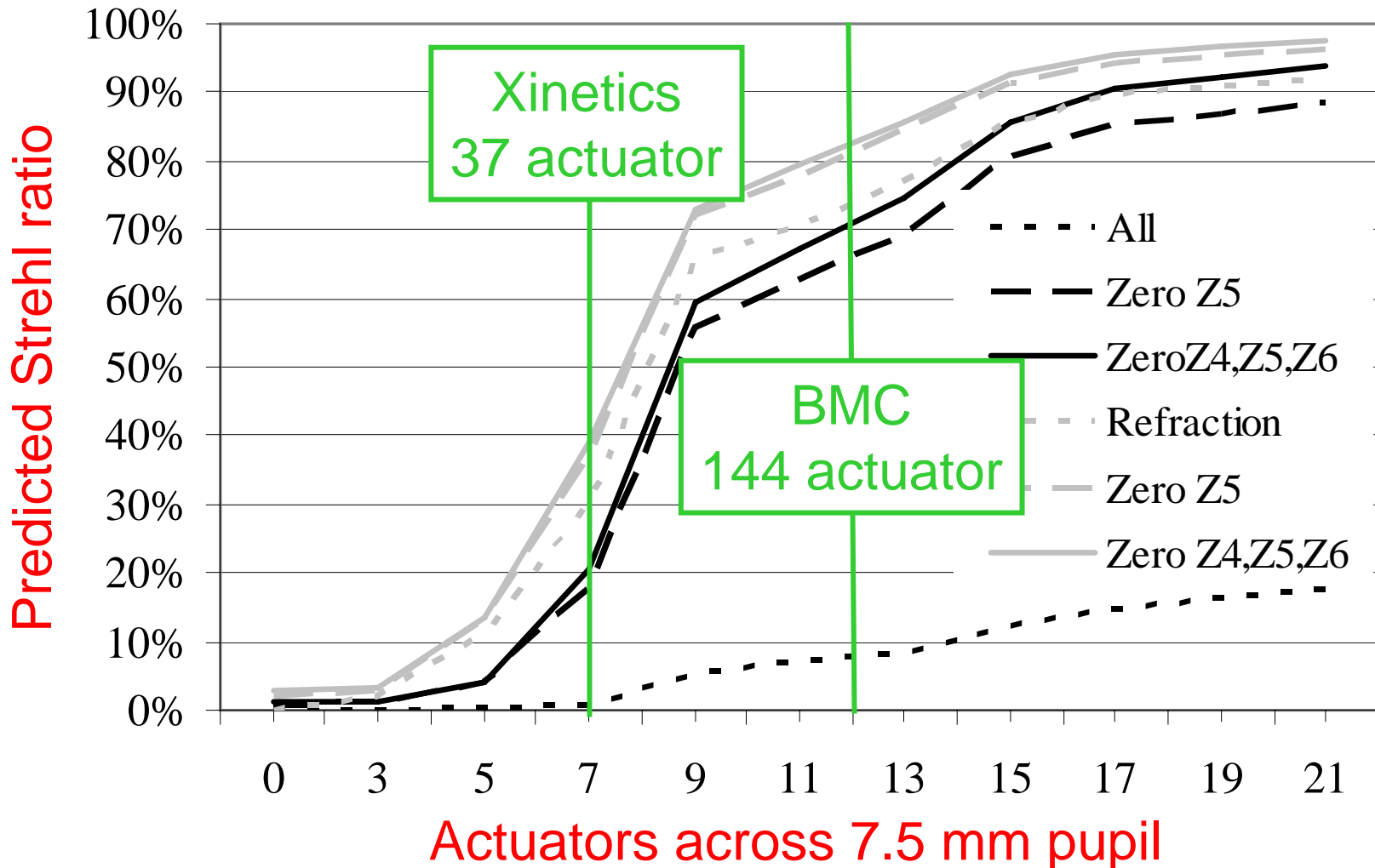
P-V wavefront error for 7.5 mm pupil



Discrete actuator deformable mirrors



Diff lim result: >14 (Roch), 11-14 (Ind)

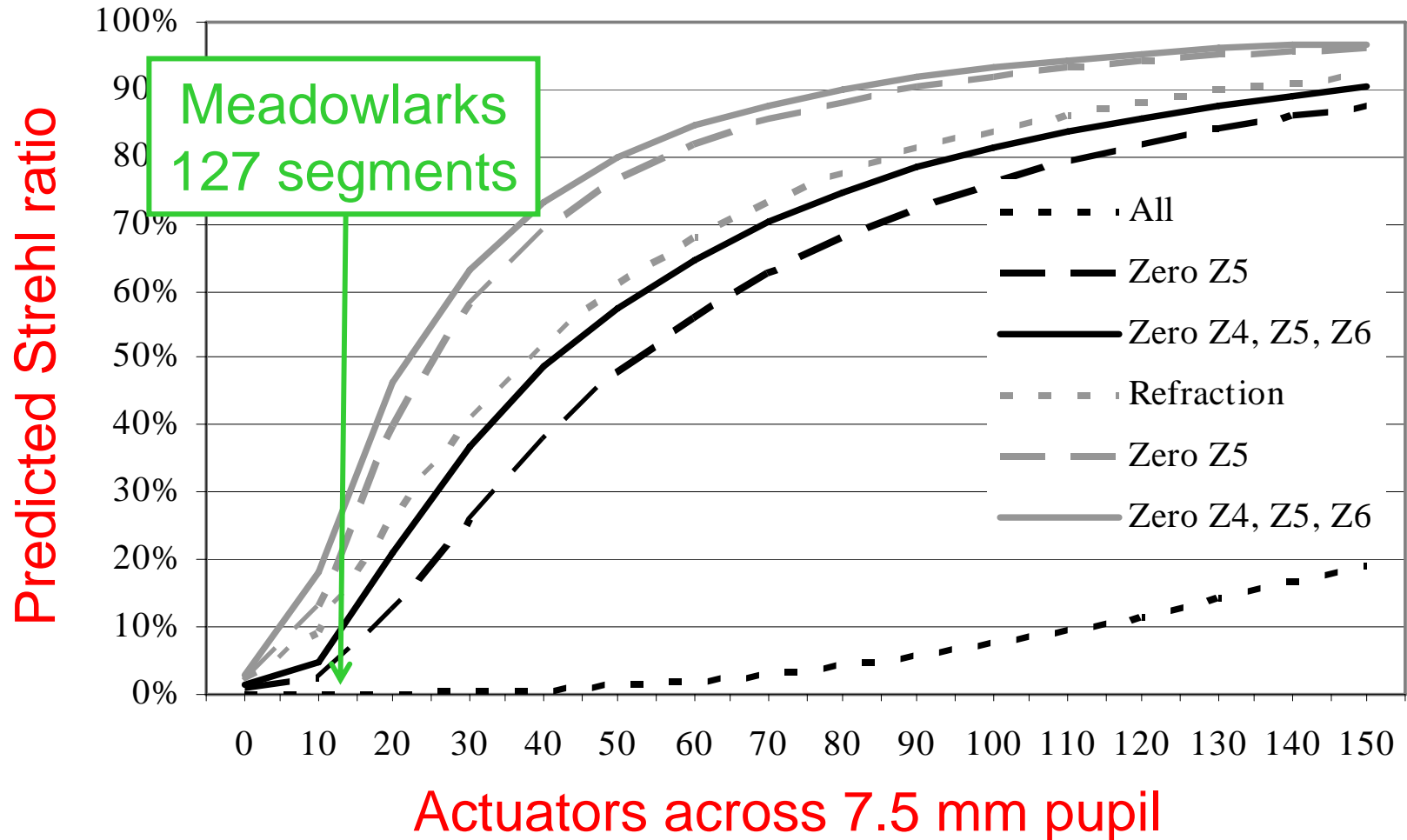


Piston segmented correctors

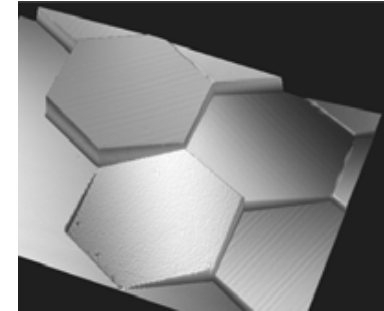


480x480 segments

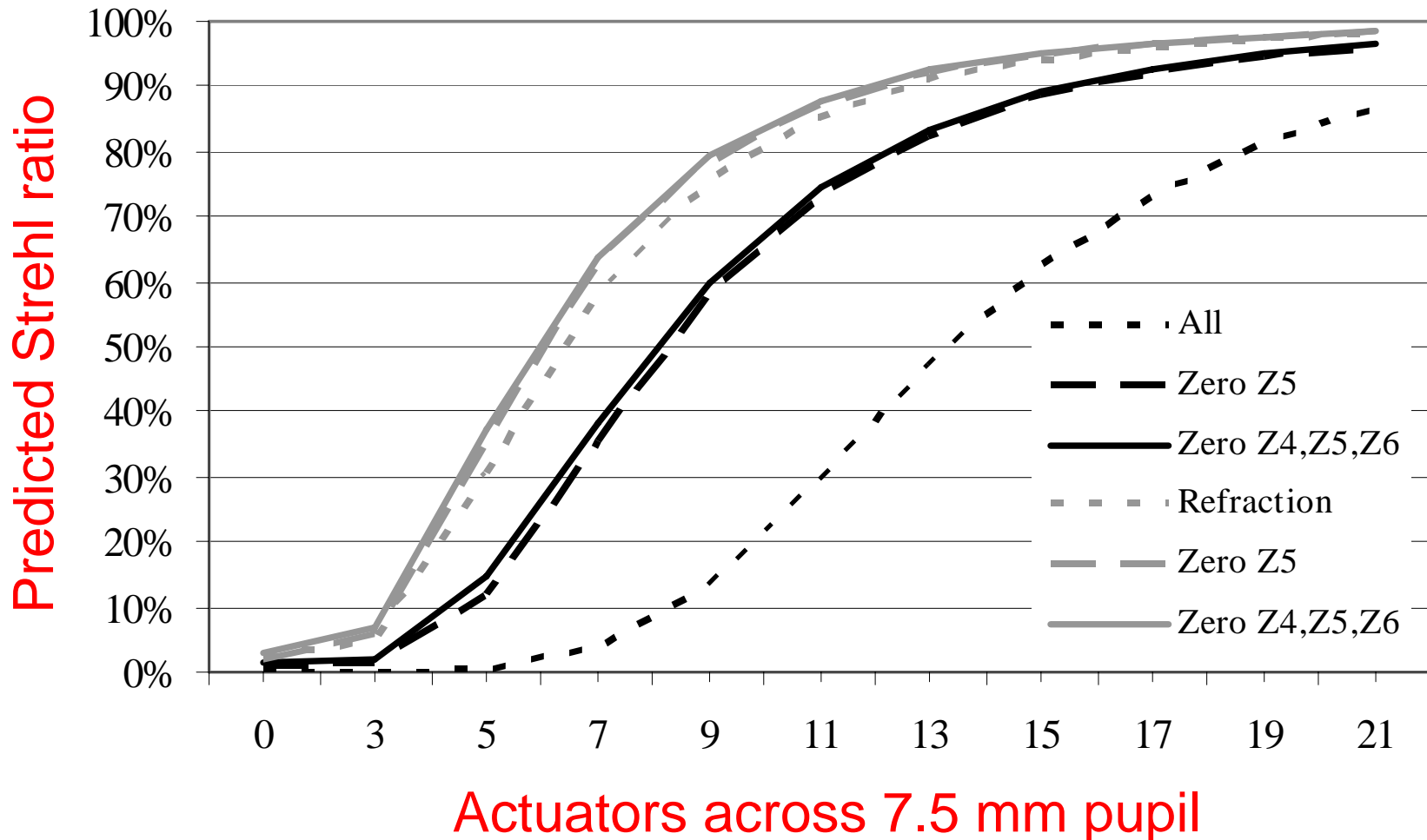
Diff lim result: >90 (Roch), 45-85 (Ind)



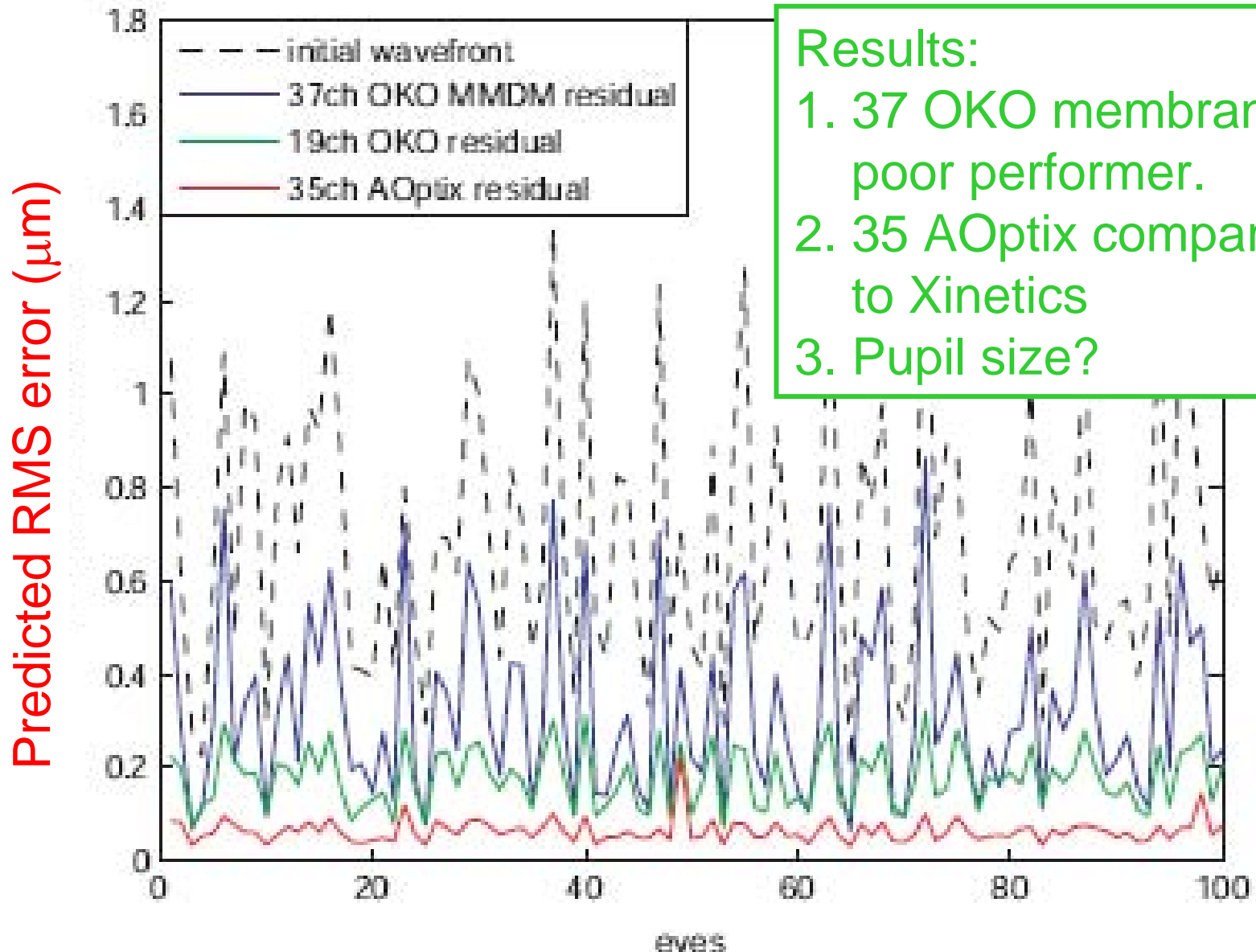
Piston/tip/tilt segmented correctors



Diff lim result: 12-19 (Roch), 9-10 (Ind)



Predicted performance for OKO and AOptix mirrors



Results:

1. 37 OKO membrane poor performer.
2. 35 AOptix comparable to Xinetics
3. Pupil size?

Desired parameters for the wavefront corrector

Temporal Bandwidth	1-6 Hz fluctuations in the eye
Reflectivity	>90% (400-900 nm)
Physical size	4 - 8 mm
Mirror Stroke (7.5-mm pupil, 95% population)	10-53 μm (Rochester) 7-11 μm (Indiana)
# of actuators or segments across 7.5 mm pupil (for 80% Strehl)	> 14 (Roch), 11-14 (Ind) Discrete actuator >> 90 (Roch), 45-85 (Ind) Piston-only segmented 12-19 (Roch), 9-10 (Ind) Piston/tip/tilt segmented >> 37 OKO membrane >19 OKO piezoelectric ~35 AOptix

Indiana adaptive optics retina camera

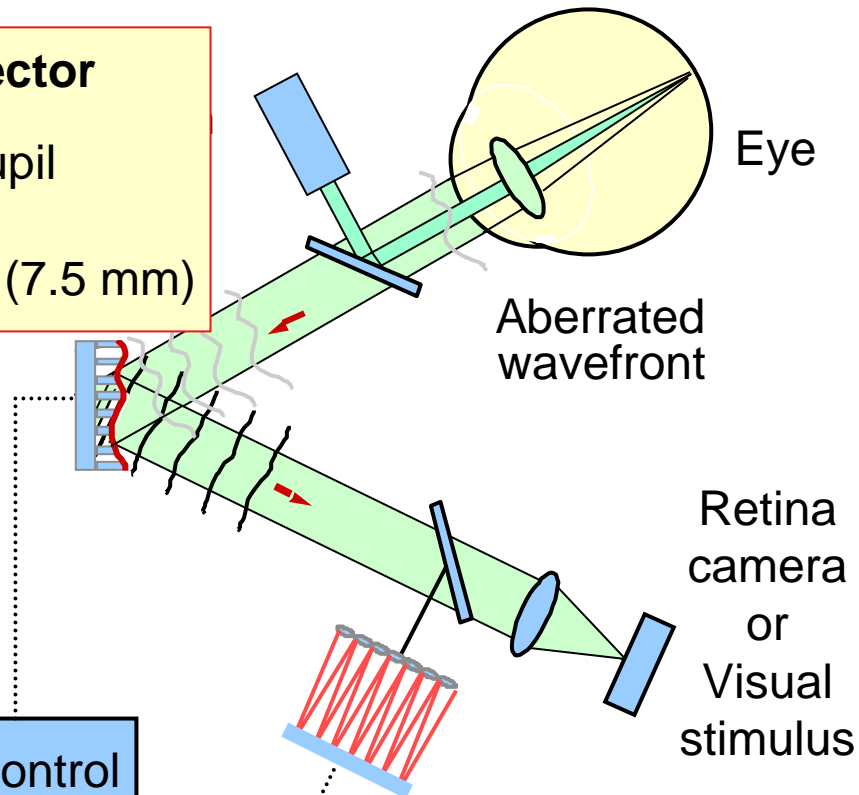
3. Xinetics wavefront corrector

37 actuators fill 6.8 mm pupil
4 μm stroke
Predicted Strehl = 0.2-0.4 (7.5 mm)

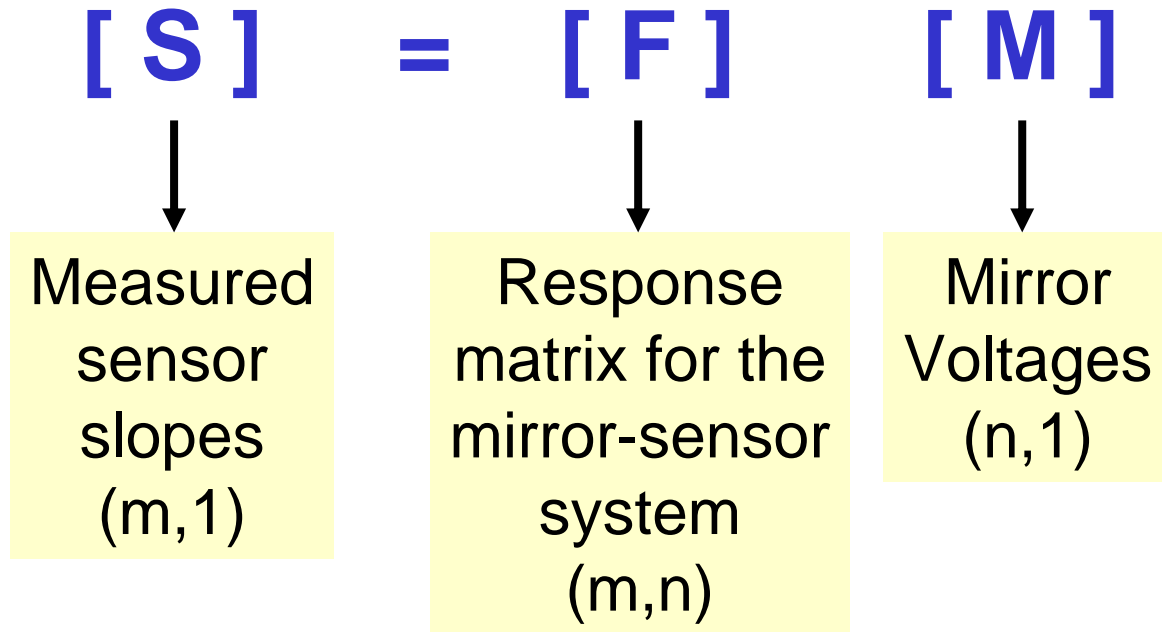
2. Control system

1. SHWS wavefront sensor

17x17 lenslets across 6.8 mm pupil
~7 pix across each focal spot
 $\lambda = 0.78 \mu\text{m}$ SLD beacon
6 μW enters eye
Inexpensive areal CCD

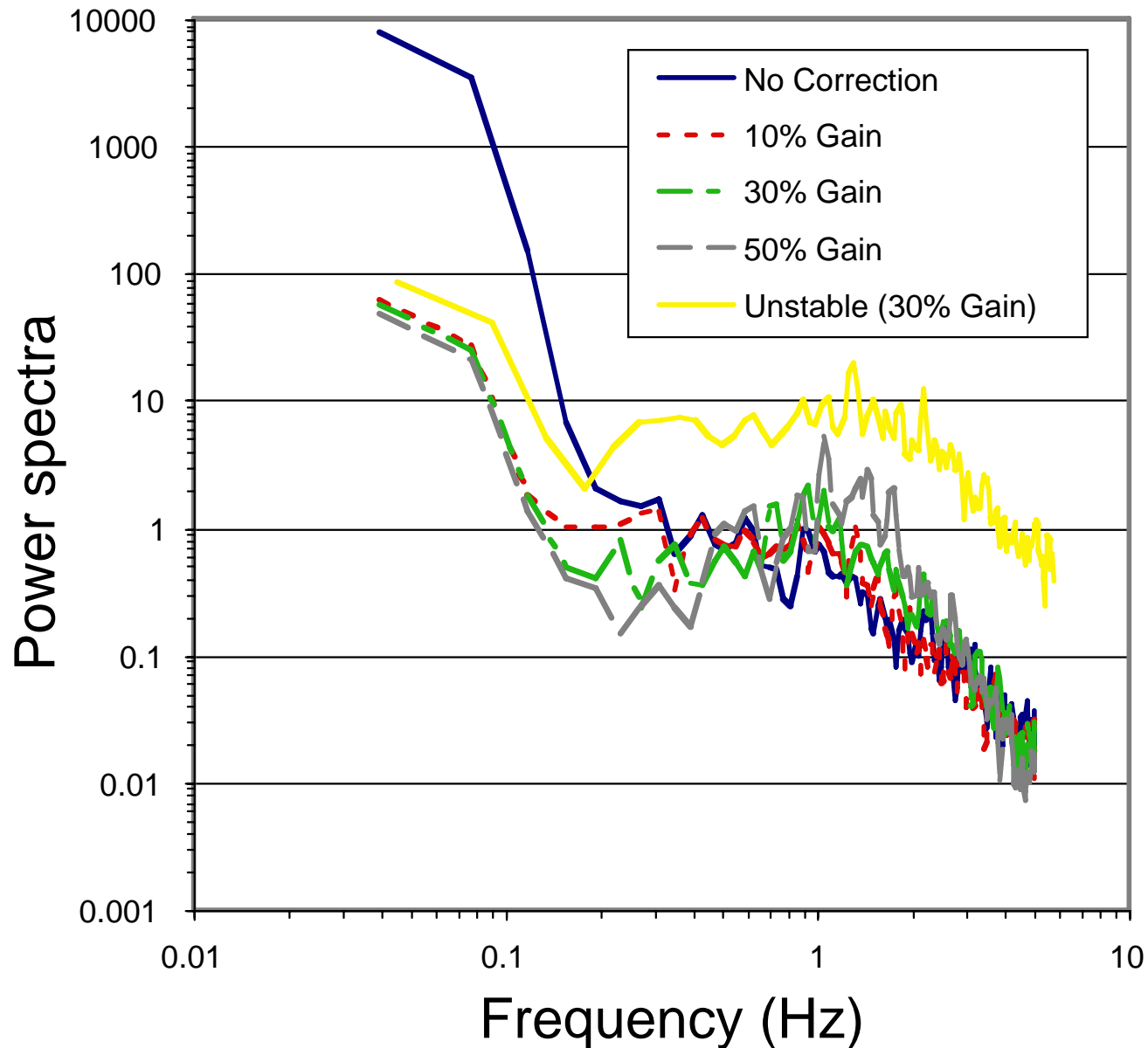


Direct slope reconstruction method

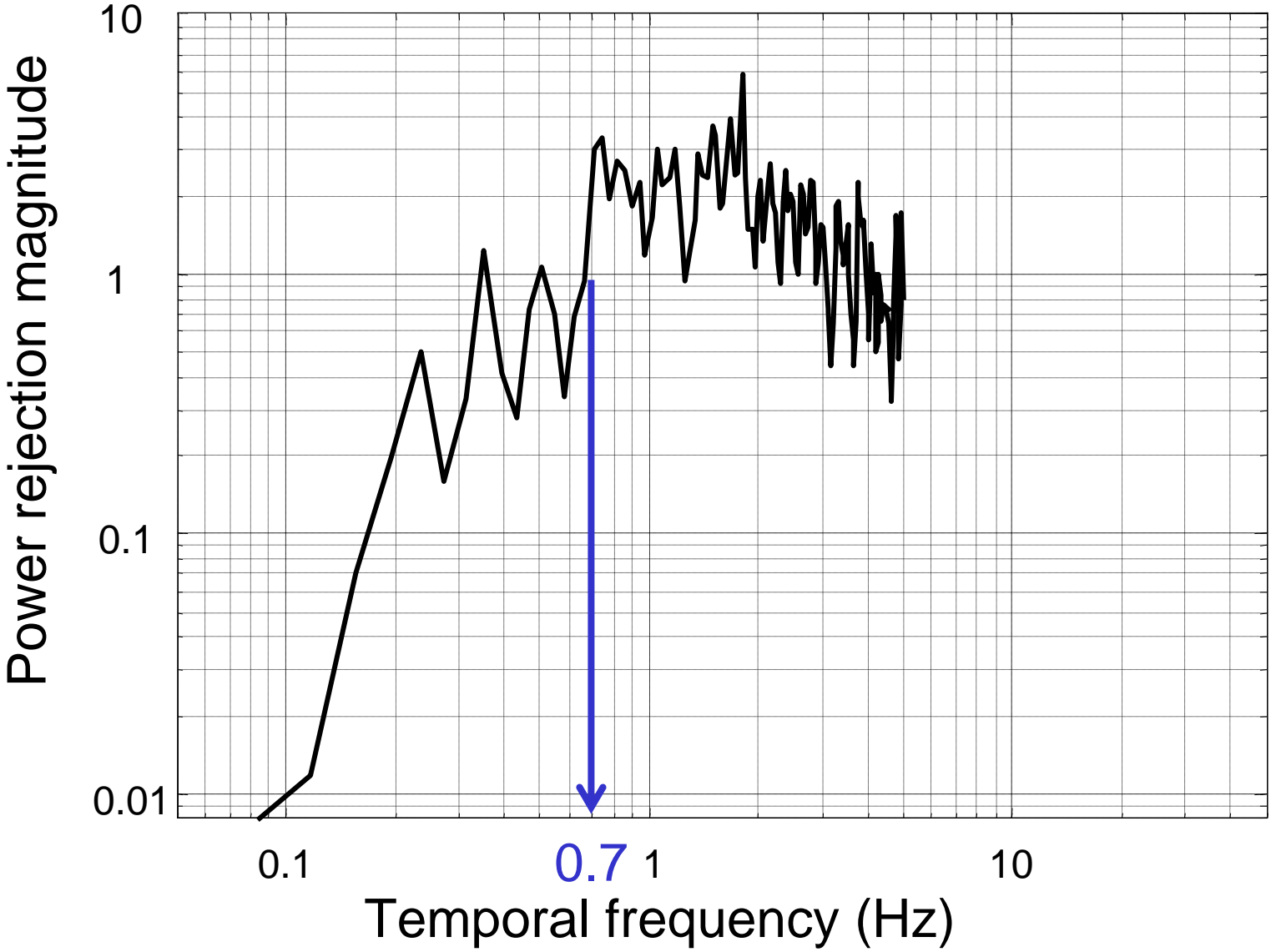


$$[M] = \underbrace{([F]^T [F])^{-1} [F]^T}_{\text{Least squares solution to inverse}} [S]$$

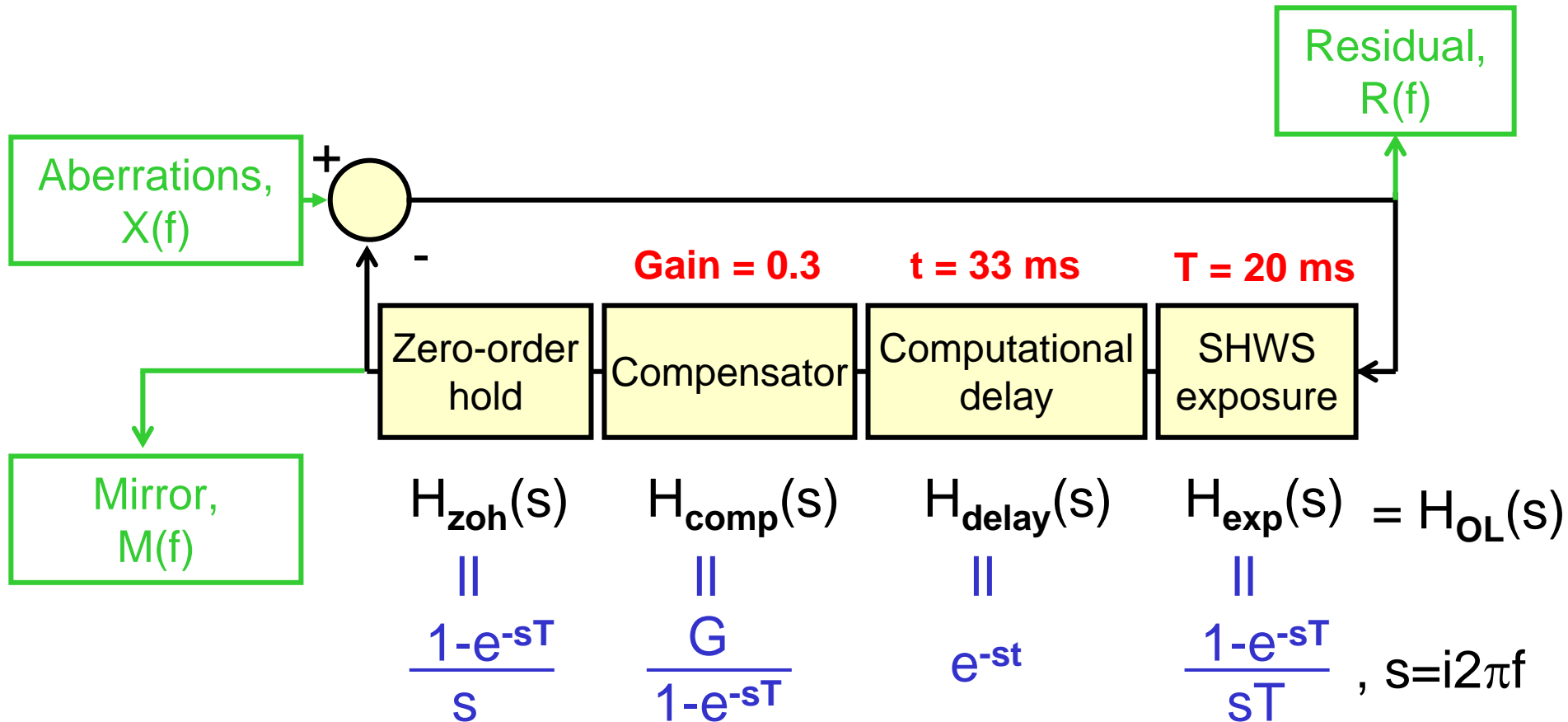
Temporal power spectra of the wavefront disturbance without and with dynamic correction



Power rejection magnitude = $\frac{\text{power spectra with AO}}{\text{power spectra w/o AO}}$

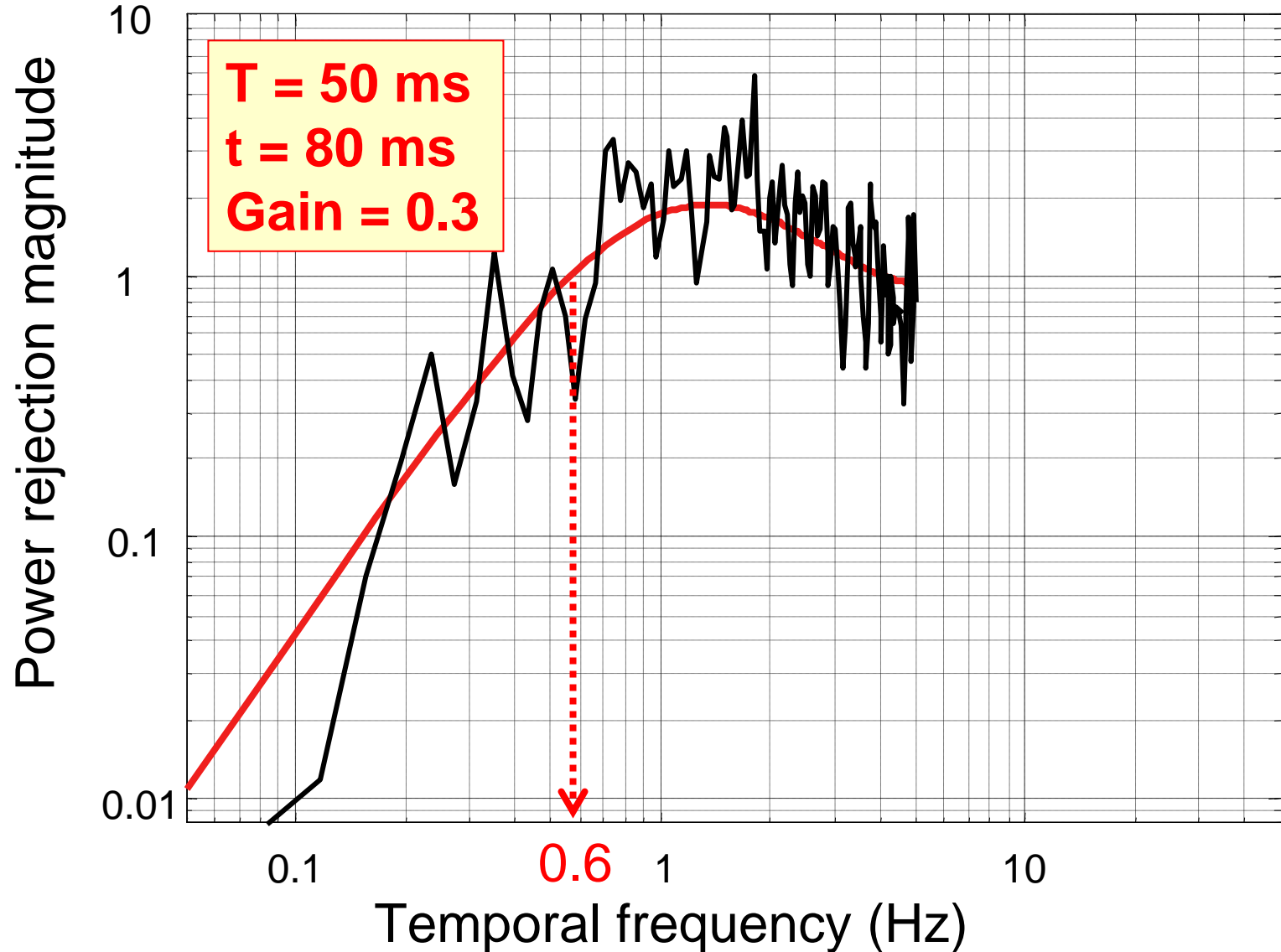


Schematic of a closed-loop AO system



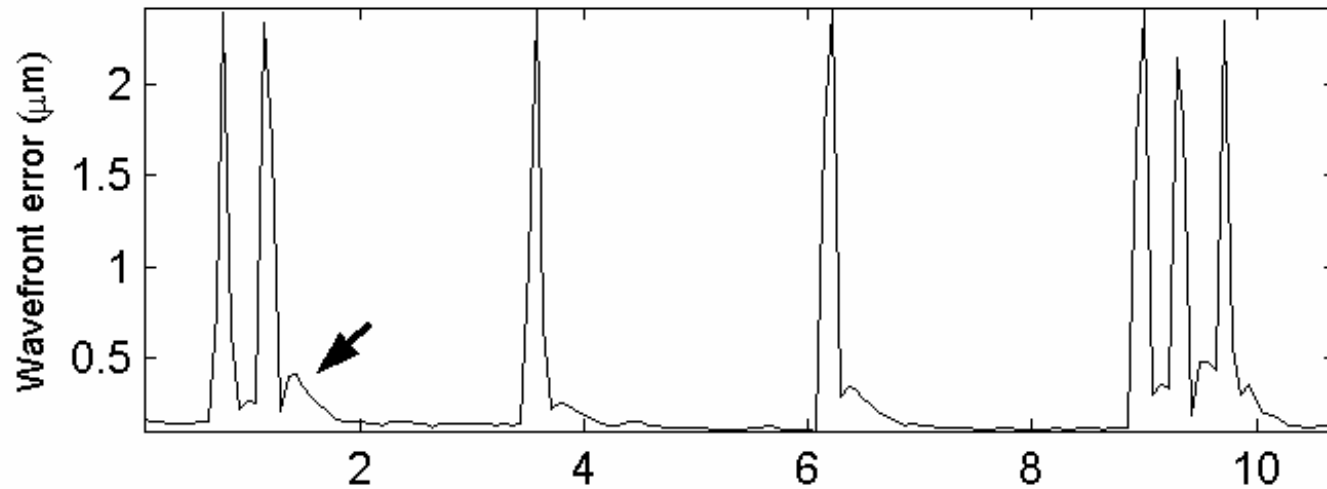
$\sqrt{\text{Power rejection curve}} = \text{Closed-loop transfer function} = H_{\text{CL}}(s) = \frac{R(s)}{X(s)} = \frac{1}{1 + H_{\text{OL}}(s)}$

Predicted and measured power rejection magnitude of the Indiana AO retina camera

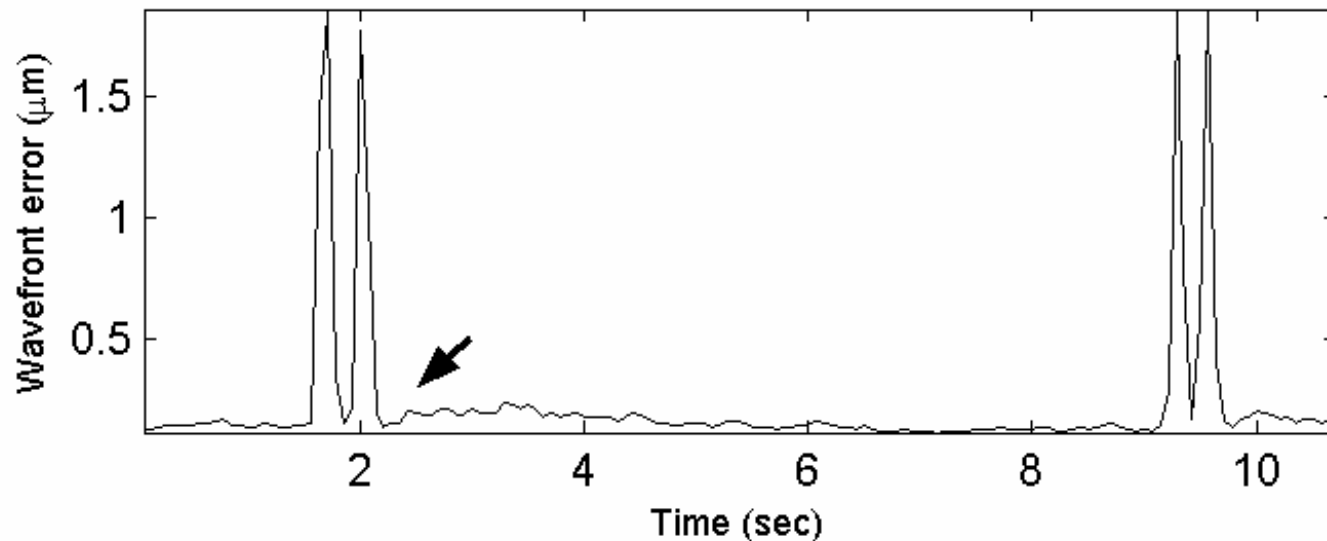


How does the AO perform during & immediately after an eye blink?

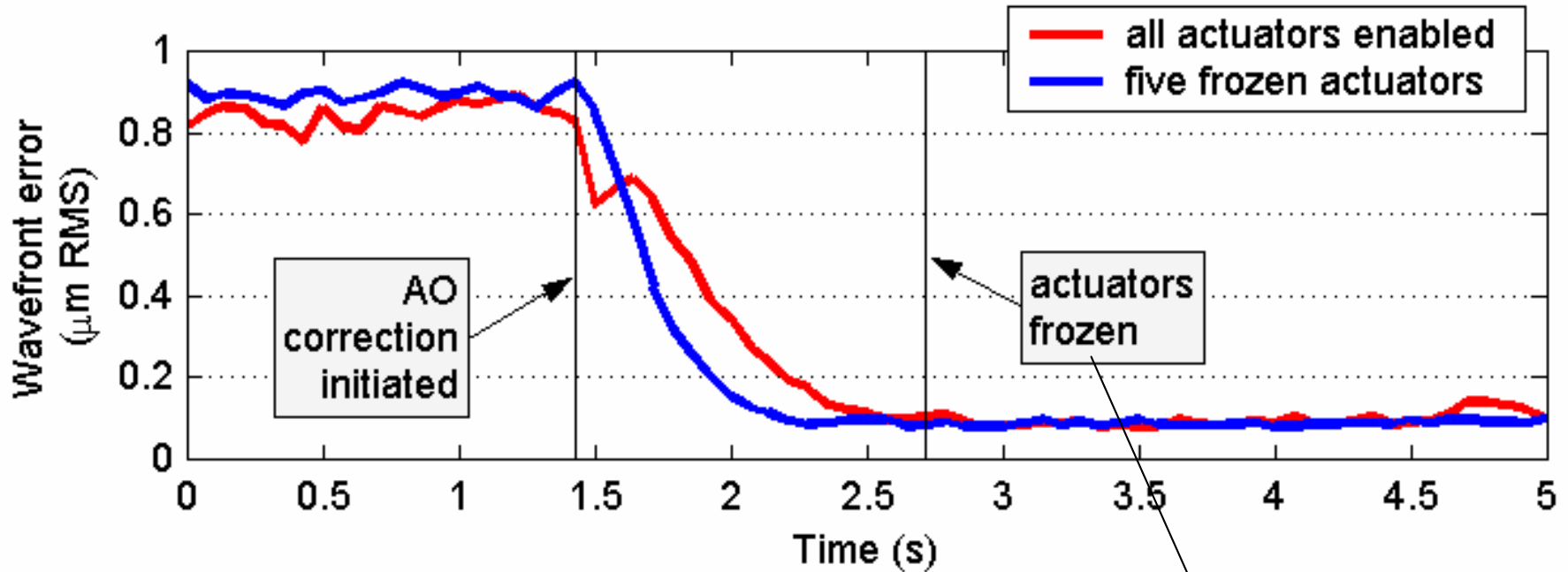
Without error suppression



With error suppression



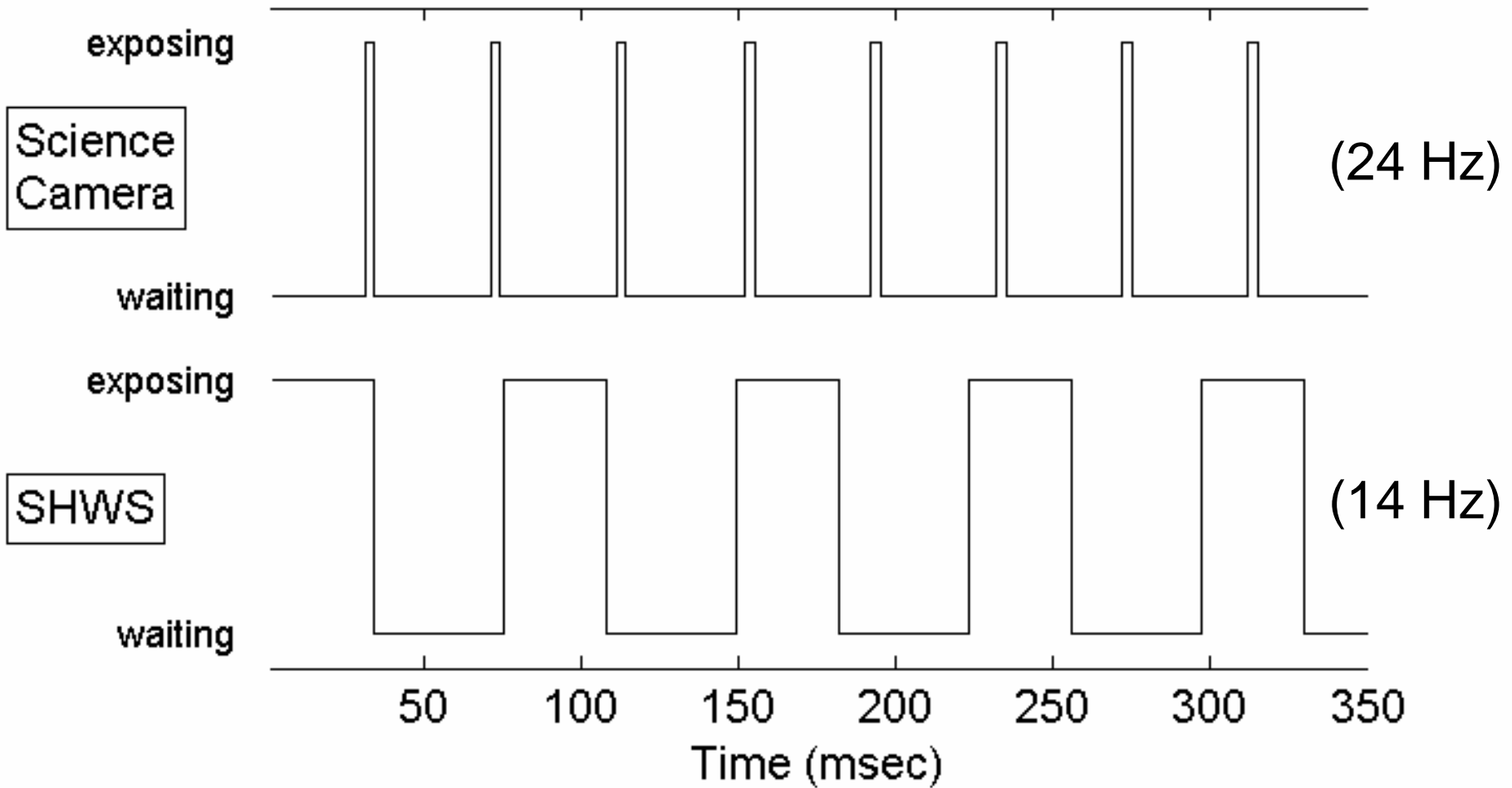
Impact of freezing the middle five actuators during a dynamic correction was found negligible



(6.8 mm pupil, $\lambda = 0.78 \mu\text{m}$)



Time stamping SHWS measurements permits more accurate deconvolution.



Timing diagram of concurrent SHWS measurements (14 Hz) and acquisition of science images (24 Hz) as obtained on one subject's eye

Indiana adaptive optics retina camera

3. Xinetics wavefront corrector

37 actuators fill 6.8 mm pupil
4 μm stroke
Predicted Strehl = 0.2-0.4 (7.5 mm)

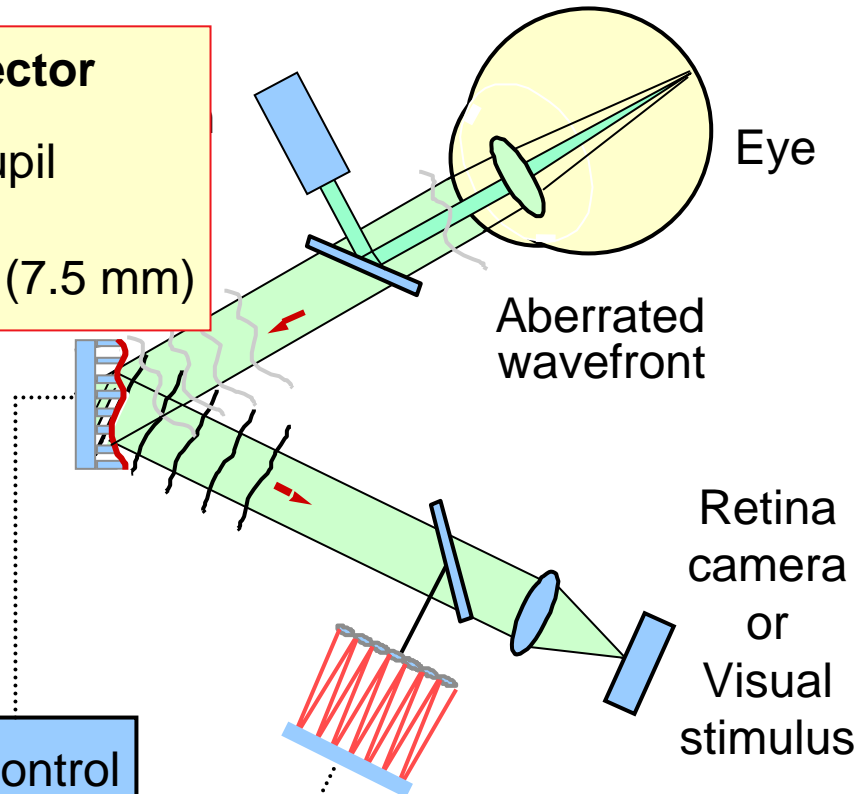
2. Control system

2. Control system

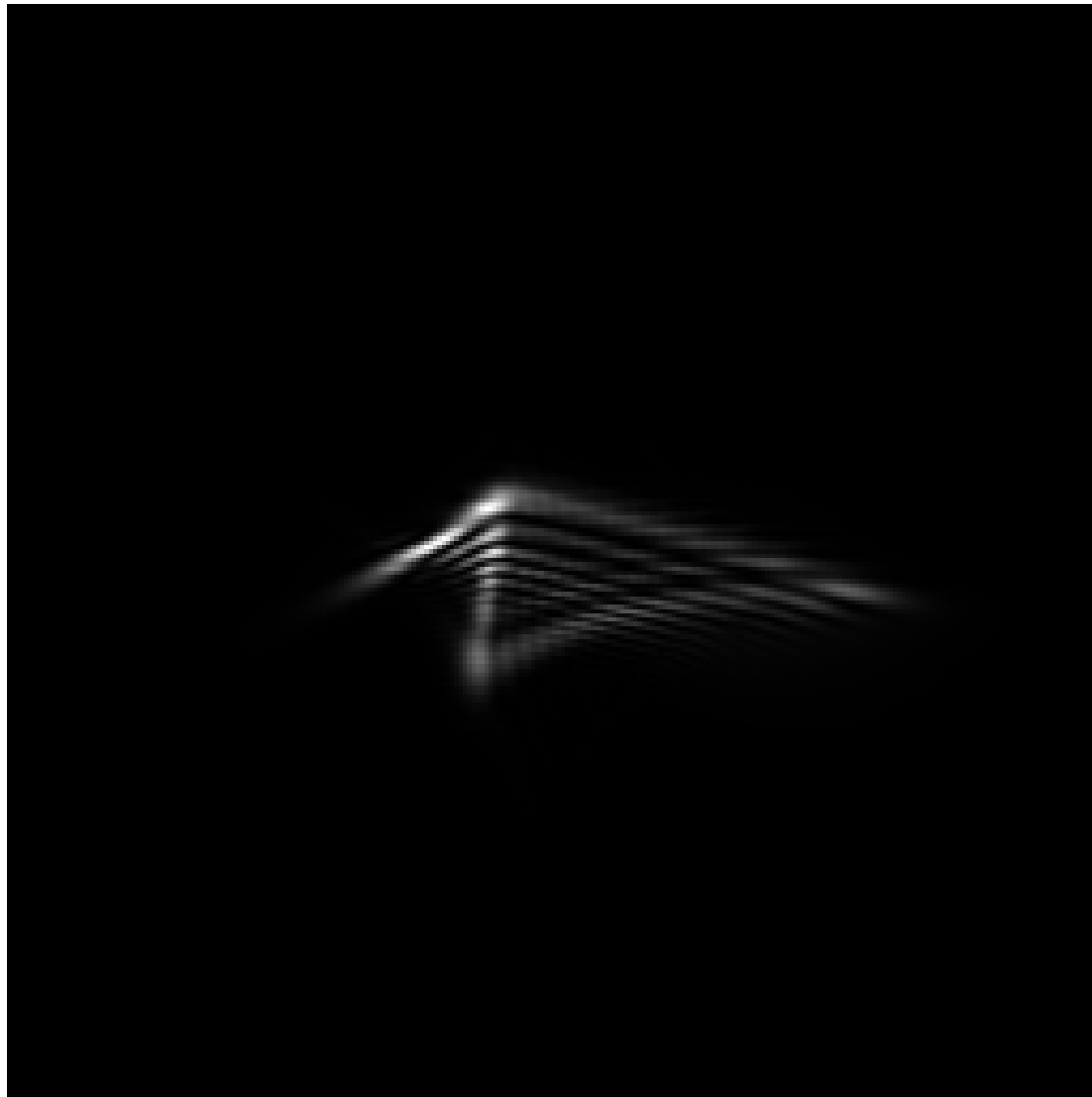
Zonal reconstructor
(direct slope)
0.6 Hz cutoff frequency

1. SHWS wavefront sensor

17x17 lenslets across 6.8 mm pupil
~7 pix across each focal spot
 $\lambda = 0.78 \mu\text{m}$ SLD beacon
6 μW enters eye
Inexpensive areal CCD



Dynamic correction in one subject's eye as revealed by measured PSF



80 frame video

- $\lambda = 0.78 \mu\text{m}$
- 6.8 mm pupil
- 10% Gain
- 21 Hz

**Before
correction:**

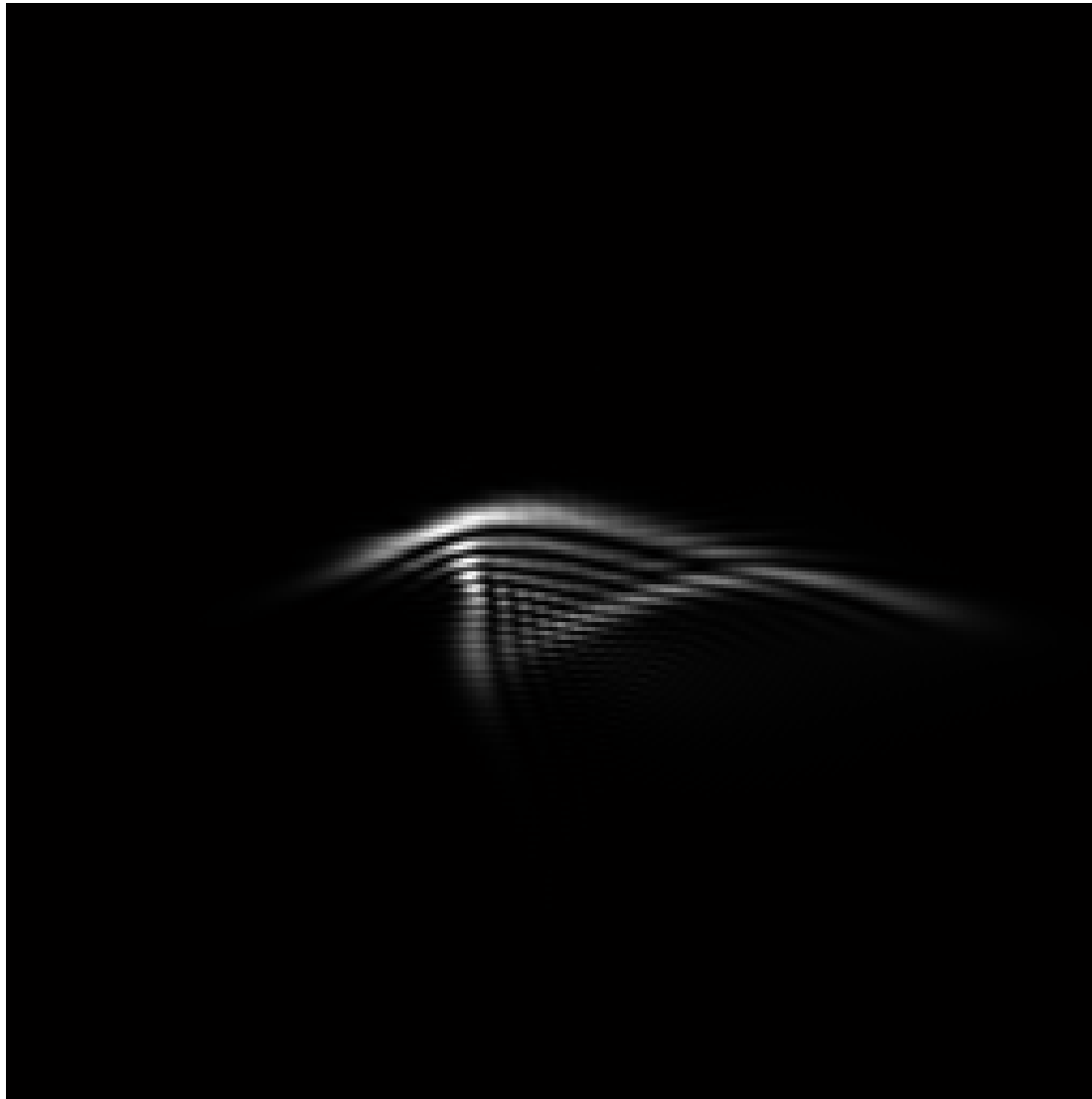
1.6 μm RMS

**After
correction:**

0.16 μm RMS

- 27 frames
to converge

Dynamic correction in one subject's eye as revealed by measured PSF



80 frame video

- $\lambda = 0.78 \mu\text{m}$
- 6.8 mm pupil
- 30% Gain
- 21 Hz

**Before
correction:**

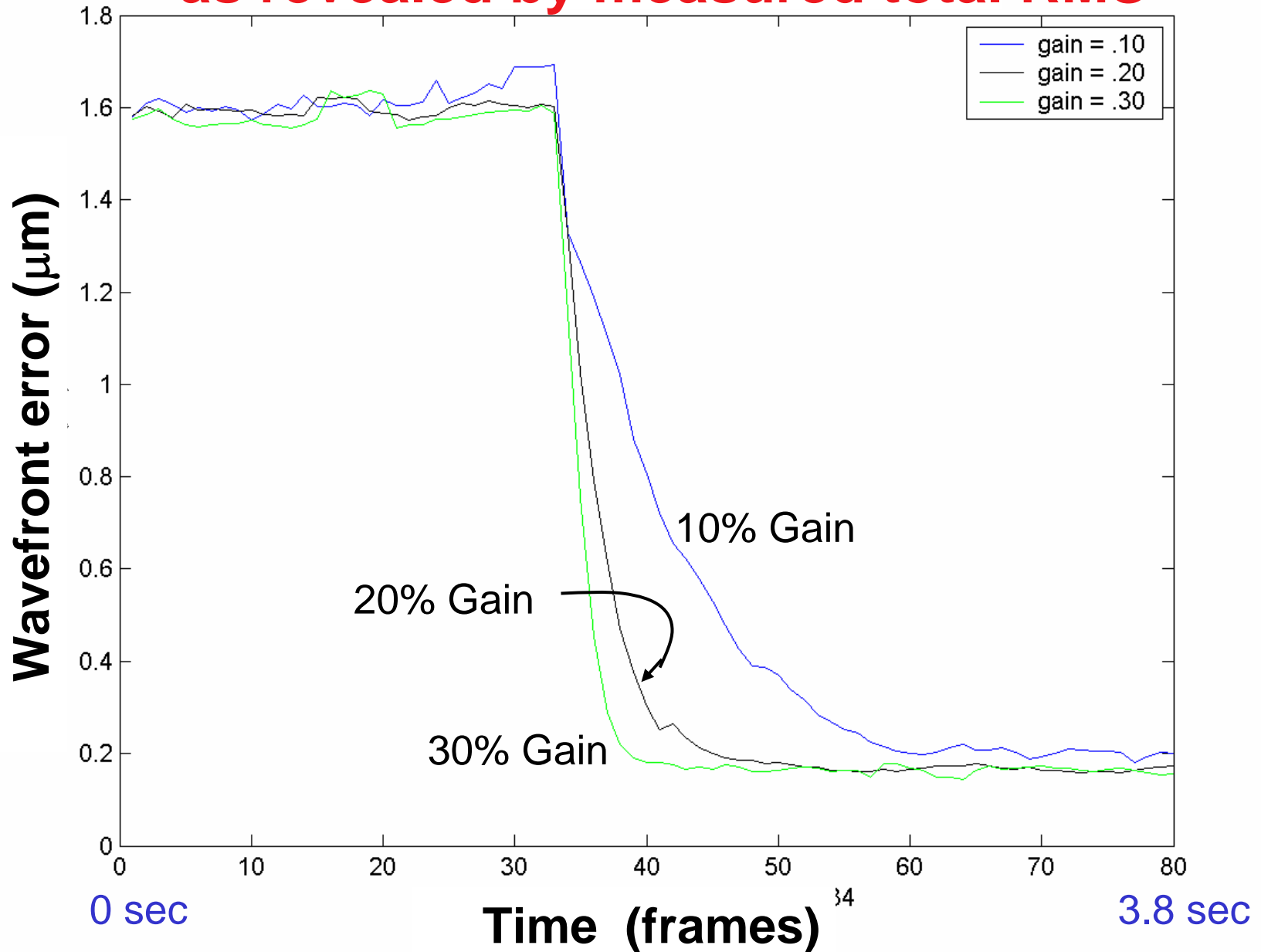
1.6 μm RMS

**After
correction:**

0.16 μm RMS

- 6 frames
to converge

Dynamic correction in one subject's eye as revealed by measured total RMS

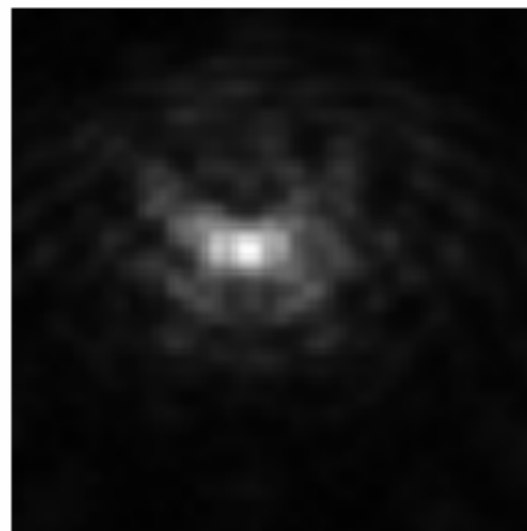
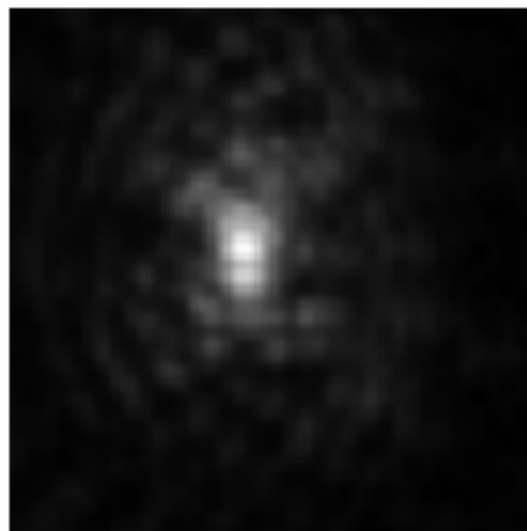


How accurate is the measured PSF?

Induced wave
aberration #1

Induced wave
aberration #2

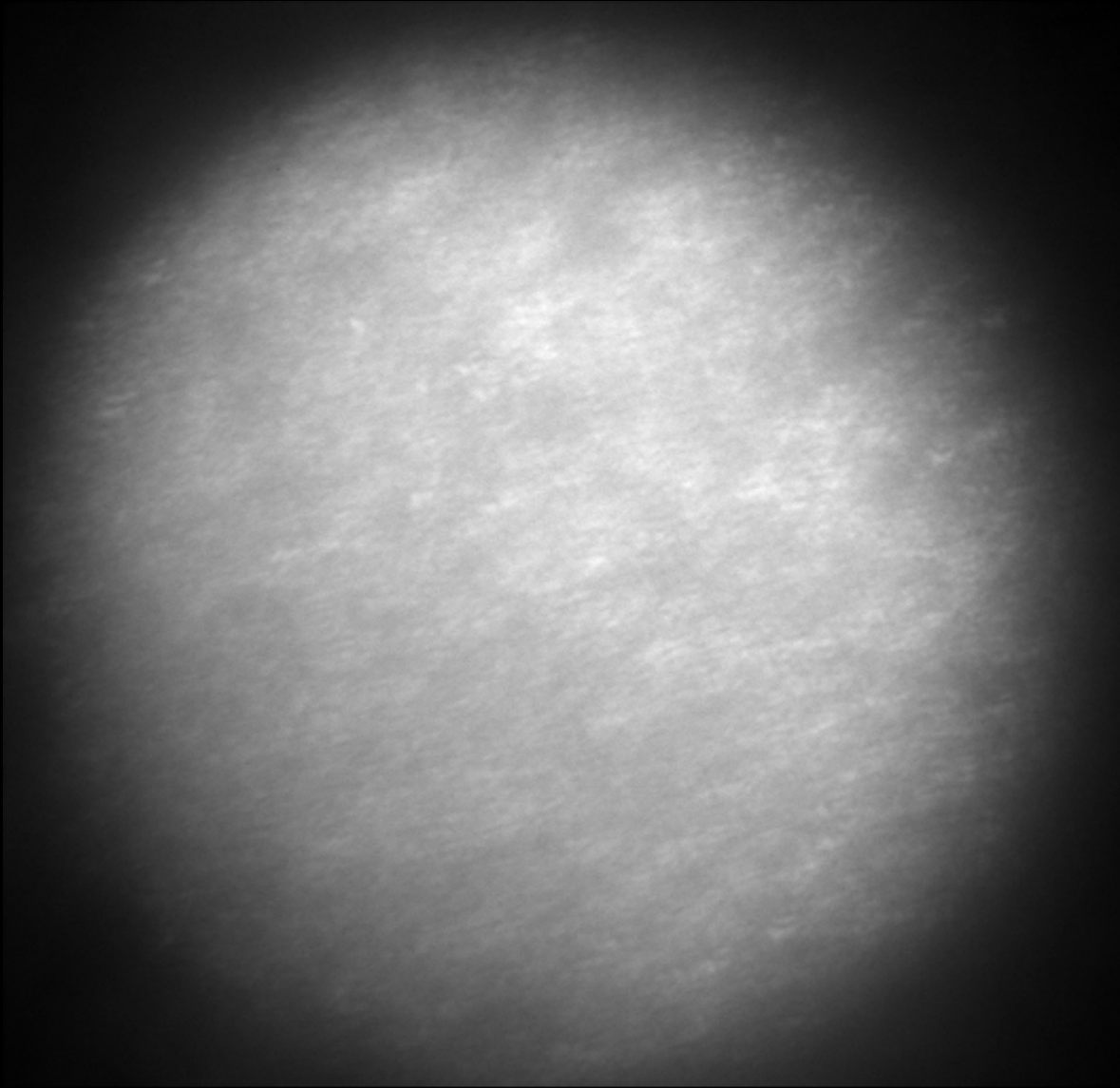
PSF recorded with
science camera (A)



PSF reconstructed
from SHWS data (B)

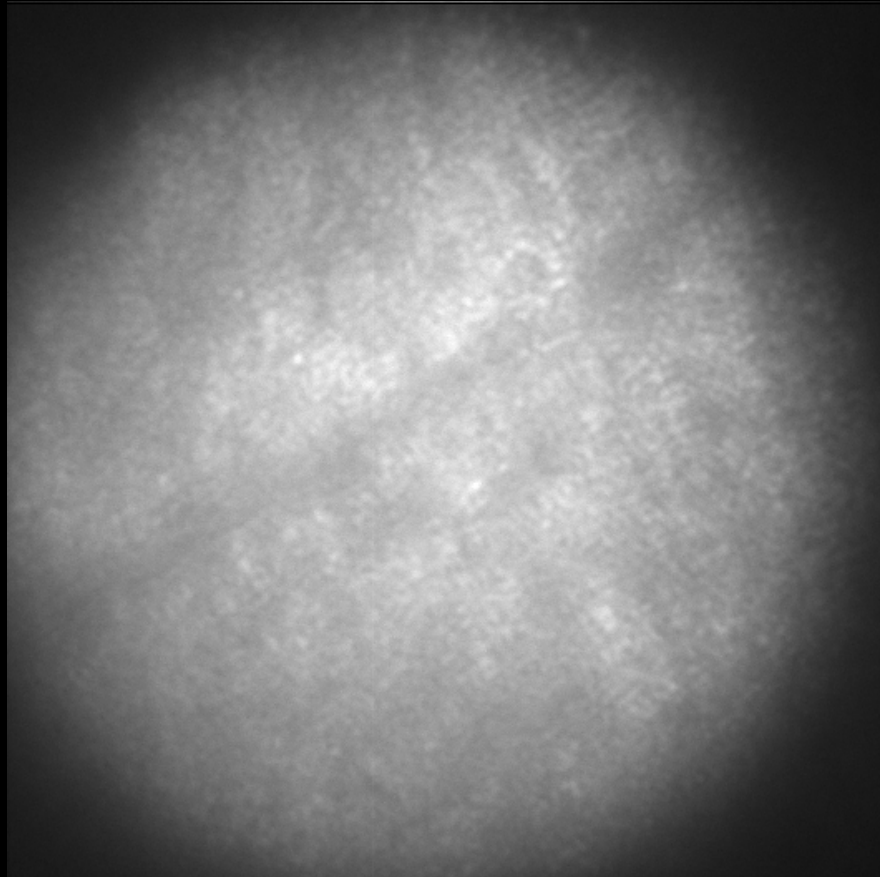


Turning on the AO during a 10 Hz video of the cone mosaic

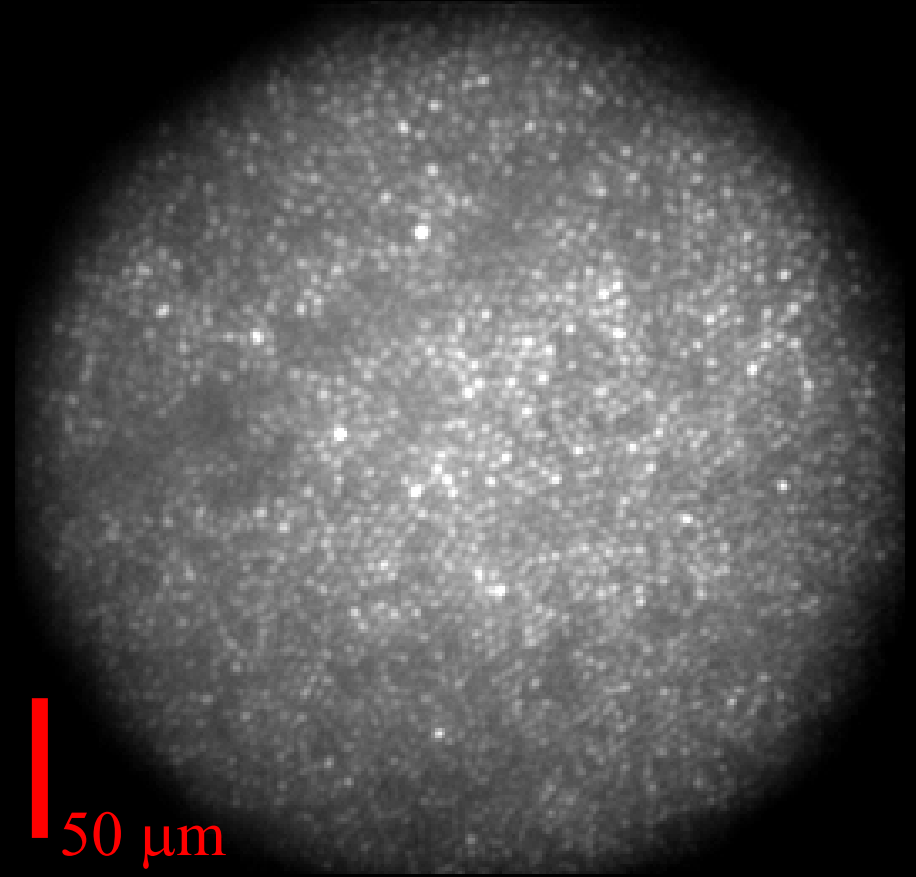


- Video rate: 10 fps
- 2° FOV
- 1° ecc.
- AO correcting at 15 Hz
- 6.0 mm pupil

Cone images in one subject's eye



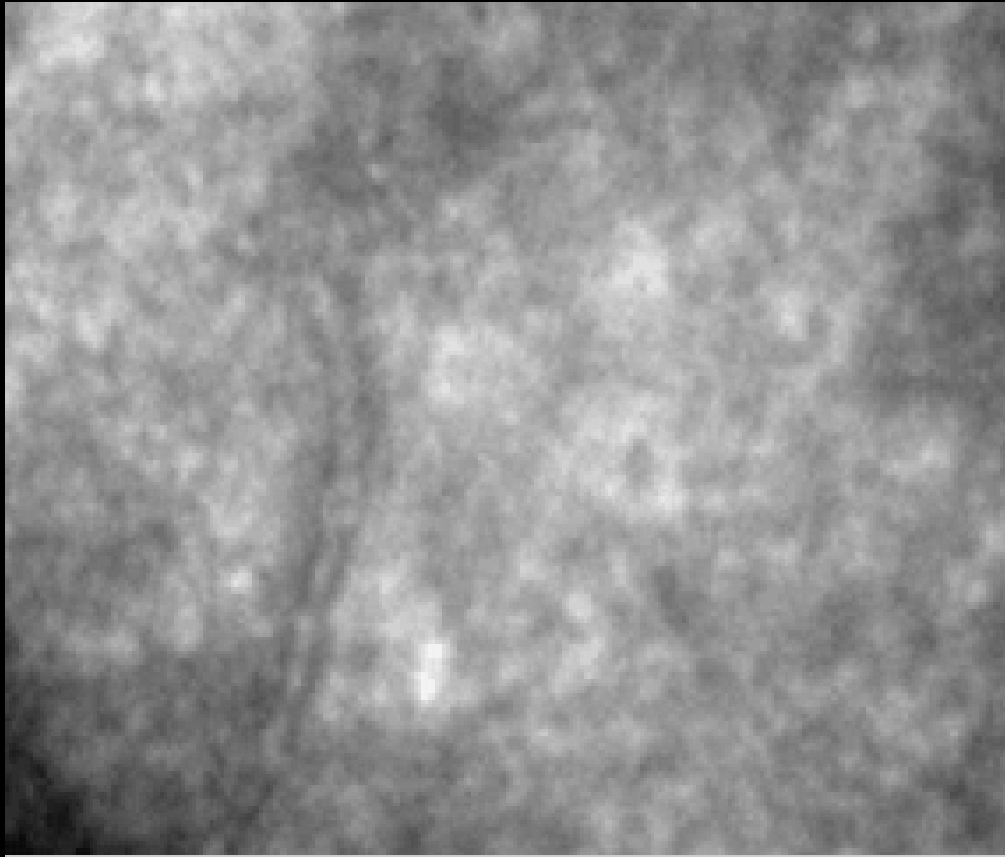
No AO correction



AO correction

1° field of view; 1.25° eccentricity

Through focus video for a fixating eye



- Video rate: ~30 fps
- $<1^\circ$ FOV
- 1.25° ecc.
- AO correcting at 22 Hz
- $0.1 \mu\text{m}$ RMS
- 6 mm pupil

RJ

Blood flow in capillaries: 4-burst imaging

Courtesy Thorn, Jonnal, Qu &
Miller - Indiana University - 2003

w/o AO

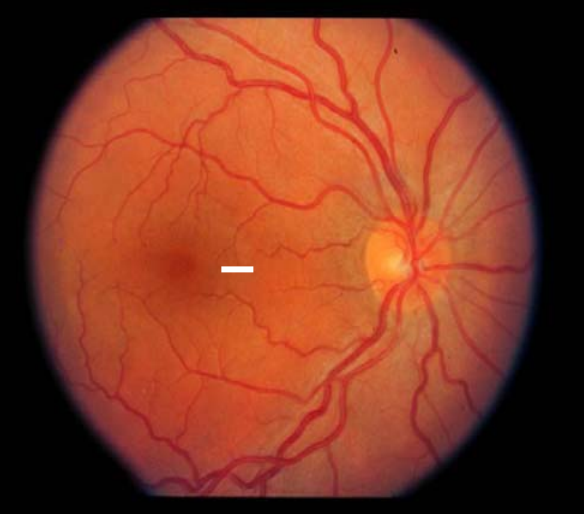
1.4 ° retinal
ecc.
 $\lambda = 679 \text{ nm}$
1 msec exp
1 msec delay

Capillary dia
~ 6.6 μm
Blood velocity
~ 1.5 mm/s

50 μm

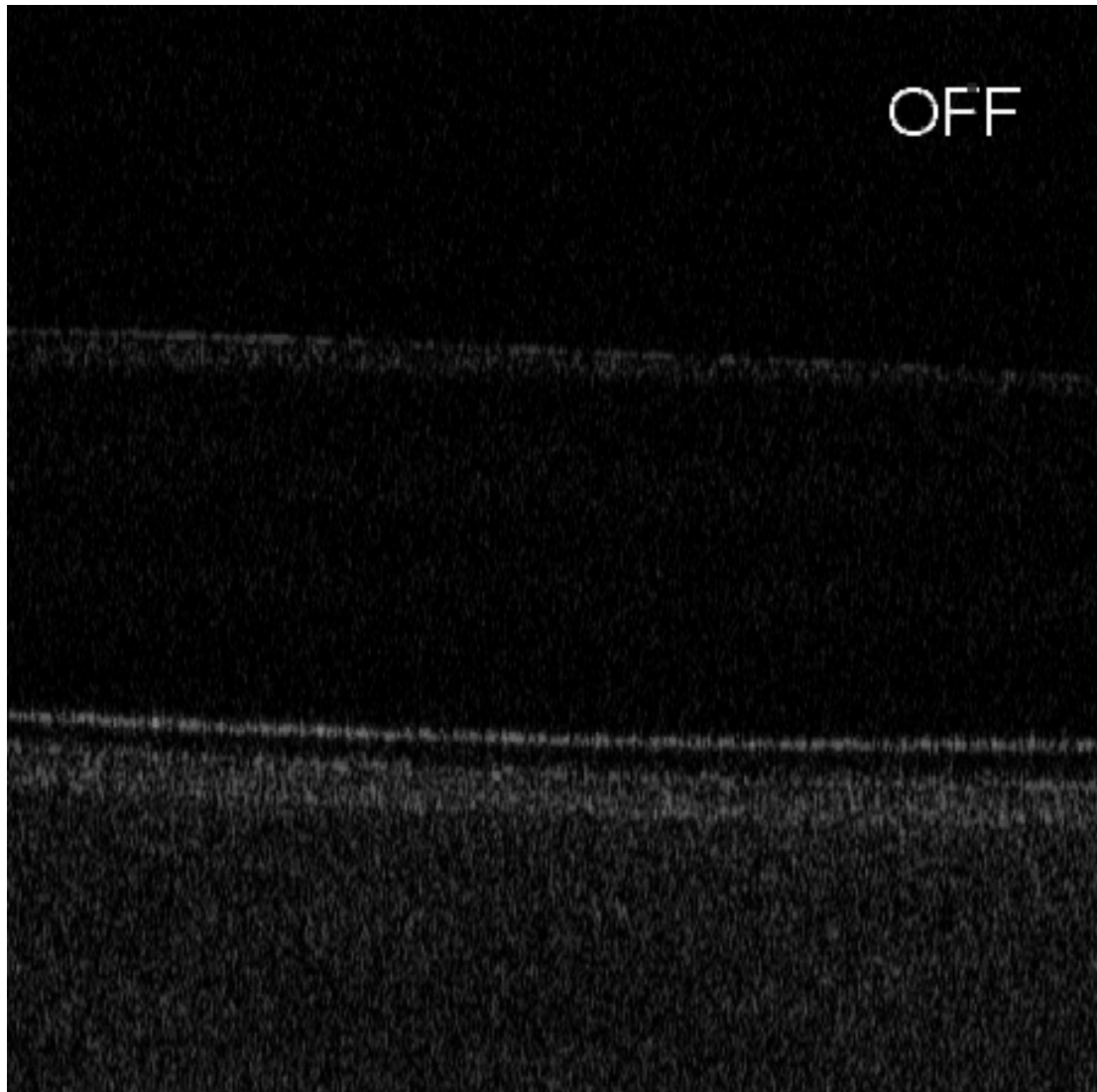
RJ

w/ AO

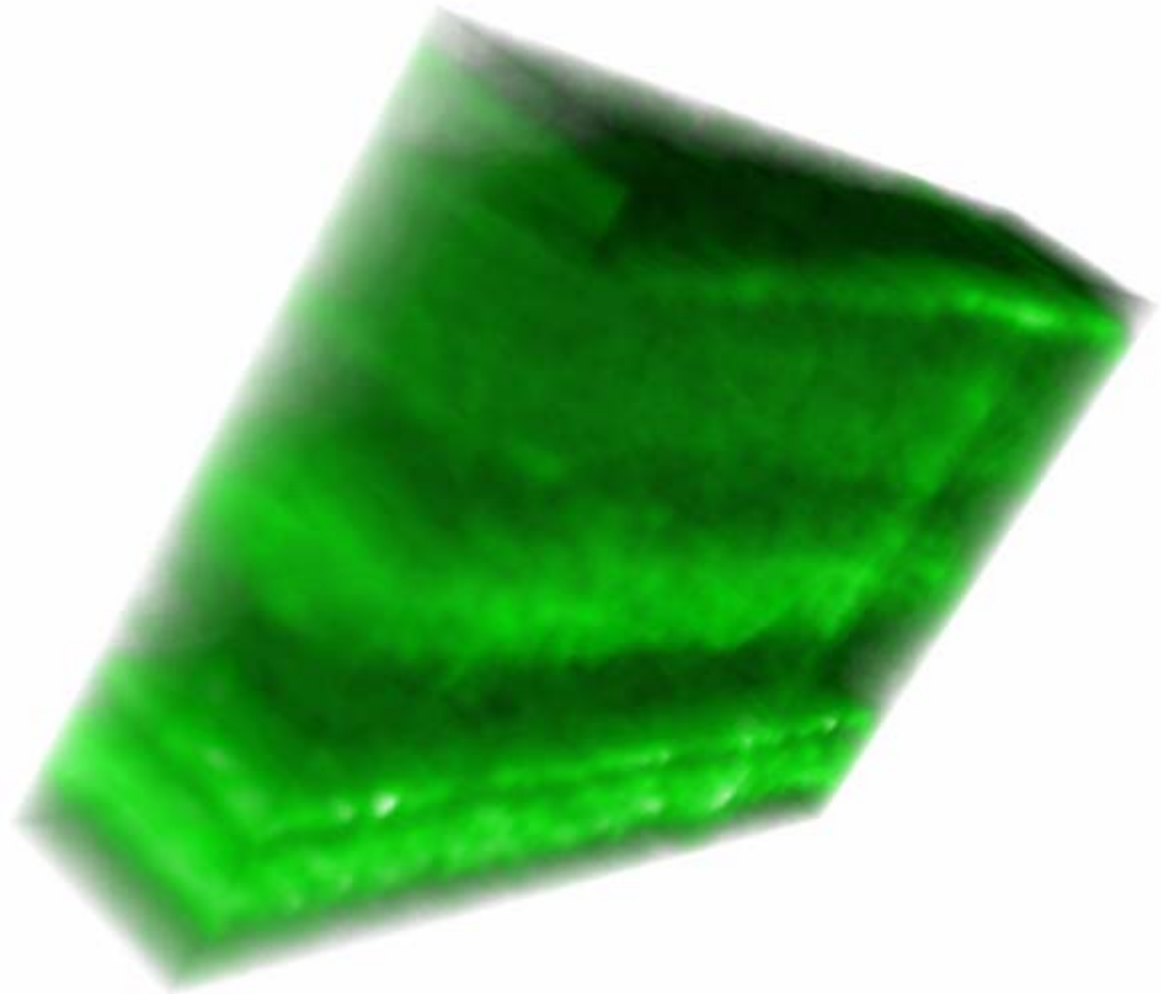
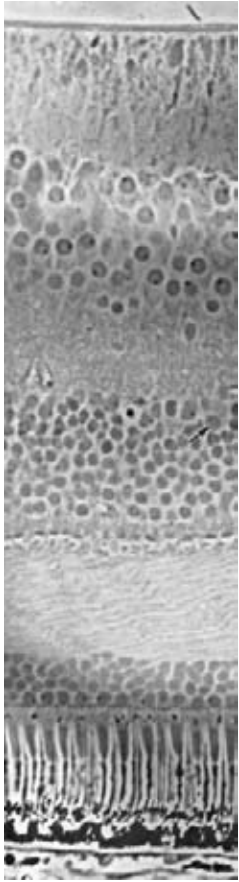


**Conventional
Retinal Image**

AO-OCT Video of Retinal Cross Section



Volumetric imaging of retina cells with AO-OCT



High 3D resolution and sensitivity permits observation of cone cells in 3 dimensions.

GROUP	TYPE	CORRECTOR	PERFORMANCE
1. Rochester	Flood cSLO	Xinetics (37,97), BMC (37) BMC (144)	< 0.1 μm RMS, d = 6.8 mm 0.12 μm RMS, d = 6 mm Coming Online
2. Houston	cSLO cSLO	Xinetics (37) BMC (144)	0.15 μm RMS, d = 7 mm
3. Indiana	Flood OCT	Xinetics (37) Xinetics (37), BMC (144) & AOptix (37)	< 0.1–0.2 μm RMS, d=6.8 mm < 0.1–0.2 μm RMS, d=6.8 mm
4. Murcia Murcia & Vienna	Flood OCT	OKO (37), LC-SLM (230,000), OKO (37)	0.1 μm RMS, d = 5.2 mm ~0.1 μm RMS, d = 5.5 mm ~0.1 μm RMS, d = 3.68 mm
5. Imperial Col / City, UK	cSLO	OKO (37)	> 0.35 μm RMS, d = 4 mm
6. Galway, Ireland	Flood	OKO (37)	
7. San Diego	Flood	OKO (19)	~0.3 μm RMS, d = 8 mm

GROUP	TYPE	CORRECTOR*	PERFORMANCE
8. UC Davis / LLNL	Flood OCT	ITEK (109) BMC (144) & Bimorph (37)	~0.1 μm RMS, d = 7 mm
9. LLNL (will be operated at Doheny)	Phoropter cSLO	BMC (144) BMC (144) and AOptix (37)	Online Online
10. Chengdu, China	Flood	Bimorph (19 or 37)	< 0.2 μm RMS, d = 5 mm
11. Paris	Flood	Bimorph (13)	< 0.3 μm RMS, d = 7 mm
12. Moscow / Kestrel Corp	Flood	Bimorph (18)	> 0.1 μm RMS, d = 4.8 mm
13. Indiana	cSLO	BMC (144)	

Summary

1. Wavefront sensor

- SHWS is used exclusively.

2. Wavefront corrector

- many types being used.
- diffraction limited imaging is very demanding.

3. Control system

- direct slope
- calibration is critical

4. Retinal imaging examples

AO does work!